Time-like and space-like electromagnetic form factors of nucleons, a unified description

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Relativistic Nuclear Physics and Quantum Chromodynamics

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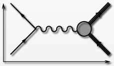
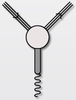
Agenda

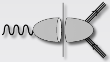


Nucleon Electromagnetic Form Factors

- Definition and properties

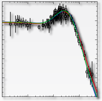


The Lomon-Gari-Krümpelmann model



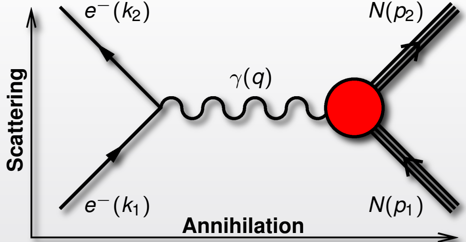
Time-like extension of the model

- Breit-Wigner regularization procedure



Data, fits, results and discussion

Dirac and Pauli Nucleon Form Factors



Scattering: $e^- N \to e^- N$

Space-like kinematic region

$$q^2 = -2\omega_1\omega_2(1-\cos\theta_e) \leq 0$$

Annihilation: $e^+e^- \leftrightarrow N\overline{N}$

Time-like kinematic region

$$q^2 = 4\omega^2 > 0$$

Scattering amplitude in Born approximation

$$\mathcal{M} = \frac{1}{q^2}\left[e\,\overline{u}(k_2)\gamma_\mu u(k_1)\right]\underbrace{\left[e\,\overline{U}(p_2)\Gamma^\mu(p_1,p_2)U(p_1)\right]}_{\text{Nucleon EM 4-current: } J_N^\mu}$$

From Lorenz and gauge invariance

$$\Gamma^\mu(p_1,p_2) = \gamma^\mu F_1^N(q^2) + \frac{i\sigma^{\mu\nu}q_\nu}{2M_N}F_2^N(q^2)$$

Dirac FF $F_1^N(q^2)$
- **Non-helicity flip**
- $F_1^N(0) = \mathcal{Q}_N$

Pauli FF $F_2^N(q^2)$
- **Helicity flip**
- $F_2^N(0) = \kappa_N$

$\mathcal{Q}_N = N$ electric charge

$\kappa_N = N$ anomalous magnetic moment

Sachs Nucleon Form Factors

Breit frame

No energy exchanged

$$p_1 = (E, -\vec{q}/2)$$
$$p_2 = (E, \vec{q}/2)$$
$$q = (0, \vec{q})$$

Nucleon EM 4-current

$$J_N^\mu = (J_N^0, \vec{J}_N) \quad \begin{cases} \rho_q = J_N^0 = e\left[F_1^N + \dfrac{q^2}{4M_N^2} F_2^N\right] \\ \vec{J}_N = e\,\overline{U}(p_2)\vec{\gamma}U(p_1)\,\left[F_1^N + F_2^N\right] \end{cases}$$

Sachs Nucleon Form Factors

$$G_M^N(q^2) = F_1^N(q^2) + F_2^N(q^2) \qquad G_E^N(q^2) = F_1^N(q^2) + \frac{q^2}{4M_N^2} F_2^N(q^2)$$

In the Breit framrepresent the **Fourier transforms** of **charge** and **magnetic moment spatial distributions** of the nucleon

Normalization at $q^2 = 0$

- $G_E^N(0) = \mathcal{Q}_N$
- $G_M^N(0) = \mu_N$

$\mu_N = \mathcal{Q}_N + \kappa_N$ is the nucleon magnetic moment

Isospin decomposition

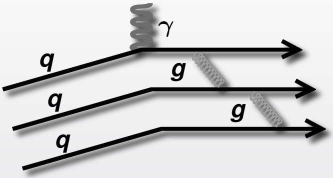
Isoscalar components

- $F_{1,2}^{is} = F_{1,2}^p + F_{1,2}^n$
- $G_{E,M}^{is} = G_{E,M}^p + G_{E,M}^n$

Isovector components

- $F_{1,2}^{iv} = F_{1,2}^p - F_{1,2}^n$
- $G_{E,M}^{iv} = G_{E,M}^p - G_{E,M}^n$

pQCD asymptotic behavior
Space-like region



- **pQCD**: as $q^2 \to -\infty$, $F_1(q^2)$ and $F_2(q^2)$ must follow counting rules
- Quarks exchange gluons to distribute momentum transferred by the photon

Dirac form factor F_1

- Non-helicity flip
- Two gluon propagators
- $F_1(q^2) \underset{q^2\to-\infty}{\sim} (-\tilde{q}^2)^{-2}$

Pauli form factor F_2

- Helicity flip
- Two gluon propagators
- $F_2(q^2) \underset{q^2\to-\infty}{\sim} (-\tilde{q}^2)^{-3}$

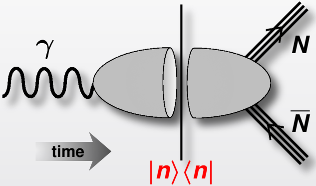
Sachs form factors G_E and G_M

- $G_{E,M}(q^2) \underset{q^2\to-\infty}{\sim} (-\tilde{q}^2)^{-2}$
- Ratio: $\dfrac{G_E}{G_M} \underset{q^2\to-\infty}{\sim}$ constant

QCD correction

$$\tilde{q}^2 = q^2 \ln\left(\frac{-q^2}{\Lambda_{\rm QCD}}\right)$$

Nucleon form factors
Time-like region ($q^2 > 0$)



- Crossing symmetry:

$$\langle N(p_2)|J^\mu|N(p_1)\rangle \to \langle \overline{N}(p_2)N(p_1)|J^\mu|0\rangle$$

- Form factors are **complex functions** of q^2

Cutkosky rule for Nucleons

$$\mathrm{Im}\langle \overline{N}(p_2)N(p_1)|J^\mu(0)|0\rangle \sim \sum_n \langle \overline{N}(p_2)N(p_1)|J^\mu(0)|n\rangle\langle n|J^\mu(0)|0\rangle \Rightarrow \begin{cases} \mathrm{Im}F_{1,2} \neq 0 \\ \text{for } q^2 > 4m_\pi^2 \end{cases}$$

$|n\rangle$ are on-shell intermediate states: 2π, 3π, 4π, ...

Time-like asymptotic behavior

Phragmèn Lindelöf theorem:

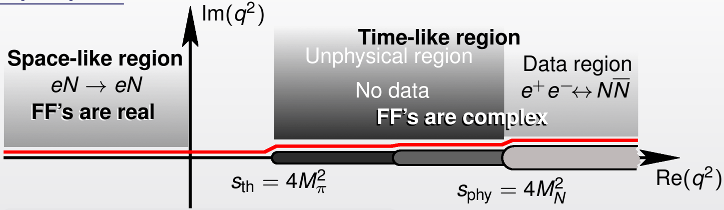
If a function $f(z) \to a$ as $z \to \infty$ along a straight line, and $f(z) \to b$ as $z \to \infty$ along another straight line, and $f(z)$ is regular and bounded in the angle between, then $a = b$ and $f(z) \to a$ uniformly in this angle.

- $$\underbrace{\lim_{q^2 \to -\infty} G_{E,M}(q^2)}_{\text{space-like}} = \underbrace{\lim_{q^2 \to +\infty} G_{E,M}(q^2)}_{\text{time-like}}$$

- $$G_{E,M} \underset{q^2 \to +\infty}{\sim} (q^2)^{-2}$$ **real**

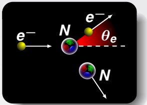
Cross sections and analyticity

q^2-complex plane



Crossing: tot. helicity $= \begin{cases} 1 \Rightarrow G_E \\ 0 \Rightarrow G_M \end{cases}$

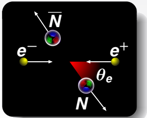
$$G_E(4M_N^2) = G_M(4M_N^2)$$



Elastic scattering

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2 \omega_2 \cos^2\frac{\theta_e}{2}}{4\omega_1^3 \sin^4\frac{\theta_e}{2}} \left[G_E^2 - \tau \left(1 + 2(1-\tau)\tan^2\frac{\theta_e}{2}\right) G_M^2 \right] \frac{1}{1-\tau}$$

$$\tau = \frac{q^2}{4M_N^2}$$



Annihilation

$$\frac{d\sigma}{d\Omega} = \frac{\alpha^2 \beta C}{4q^2} \left[(1+\cos^2\theta)|G_M|^2 + \frac{1}{\tau}\sin^2\theta |G_E|^2 \right]$$

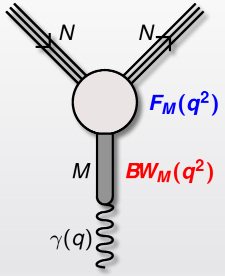
$$\beta = \sqrt{1 - \frac{1}{\tau}}$$

The Lomon-Gari-Krünpelmann Model

The Model

The model is based on an idea of Iachello, Jackson and Landé, later developed by Gari and Krümpelmann, and Lomon. Nucleon EMFFs are parametrized using a **mixture** of Vector Meson Dominance (**VMD**) and **pQCD**.

- At low energy the coupling to the photons is described through vector meson exchange [VMD in terms of **propagators** $BW_M(q^2)$, $M = \rho, \omega, \phi, \rho', \omega'$].
- **Hadron/quark form factors** $F_M(q^2)$ at vector meson-nucleon (quark) vertices control transition to perturbative QCD at high momentum transfer.



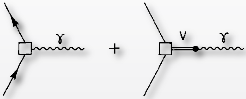
Analytic extension: space-like $\longrightarrow$ time-like

- **$BW_M(q^2)$ for broad mesons**:
  simple poles $\longrightarrow$ poles with finite energy-dependent widths
- **Dispersion relations**:
  rigorous analytic continuation of $BW_M(q^2)$ from time-like to space-like region

Parametrization for the isospin components

The **VMD** parametrization is for isospin components to single out different species of vector meson contributions

Isoscalar: ω, ω', ϕ Isovector: ρ, ρ'



Isospin components of Dirac and Pauli EMFFs

$$F_1^{iv}(q^2) = \left[BW_\rho(q^2) + BW_{\rho'}(q^2)\right] F_1^\rho(q^2) + \left[1 - BW_\rho(0) - BW_{\rho'}(0)\right] F_1^D(q^2)$$

$$F_2^{iv}(q^2) = \left[\kappa_\rho BW_\rho(q^2) + \kappa_{\rho'} BW_{\rho'}(q^2)\right] F_2^\rho(q^2) + \left[\kappa_{iv} - \kappa_\rho BW_\rho(0) - \kappa_{\rho'} BW_{\rho'}(0)\right] F_2^D(q^2)$$

$$F_1^{is}(q^2) = \left[BW_\omega(q^2) + BW_{\omega'}(q^2)\right] F_1^\omega(q^2) + BW_\phi(q^2) F_1^\phi(q^2) + [1 - BW_\omega(0) - BW_{\omega'}(0)] F_1^D(q^2)$$

$$F_2^{is}(q^2) = \left[\kappa_\omega BW_\omega(q^2) + \kappa_{\omega'} BW_{\omega'}(q^2)\right] F_2^\omega(q^2) + \kappa_\phi BW_\phi(q^2) F_2^\phi(q^2) + [\kappa_{is} - \kappa_\omega BW_\omega(0) - \kappa_{\omega'} BW_{\omega'}(0)] F_2^D(q^2)$$

The ingredients...

$$F_i^{\mathrm{I}}(q^2) = \sum_{V_{\mathrm{I}}} \kappa_{V_{\mathrm{I}}}^{i-1} BW_{V_{\mathrm{I}}}(q^2) F_i^{V_{\mathrm{I}}}(q^2) + \left[\kappa_{\mathrm{I}}^{i-1} - \sum_{V_{\mathrm{I}}} \kappa_{V_{\mathrm{I}}}^{i-1} BW_{V_{\mathrm{I}}}(0) F_i^{V_{\mathrm{I}}}(0)\right] F_i^D(q^2)$$

$BW_{V_{\mathrm{I}}}(q^2)$

$$BW_{V_{\mathrm{I}}}(q^2) = \frac{g_{V_{\mathrm{I}}} M_{V_{\mathrm{I}}}^2}{f_{V_{\mathrm{I}}}} \times [V_{\mathrm{I}}\text{-propagator}] \begin{cases} (M_{V_{\mathrm{I}}}^2 - q^2)^{-1} & V_{\mathrm{I}} = \omega, \phi \\ [\text{analytic}] & V_{\mathrm{I}} = \rho, \rho', \omega' \end{cases}$$

Meson nucleon FFs

$V_I = \rho, \rho', \omega, \omega'$

$$F_i^{V_{\mathrm{I}}}(q^2) = f_i(q^2) = \left(\frac{\Lambda_1^2}{\Lambda_1^2 - \widetilde{q}^2}\right)\left(\frac{\Lambda_2^2}{\Lambda_2^2 - \widetilde{q}^2}\right)^i$$

$$F_1^{\phi}(q^2) = f_1(q^2)\left(\frac{\widetilde{q}^2}{\widetilde{q}^2 - \Lambda_1^2}\right)^{3/2}$$

$$F_2^{\phi}(q^2) = f_2(q^2)\left(\frac{\Lambda_1^2}{\mu_\phi^2}\frac{\widetilde{q}^2 - \mu_\phi^2}{\widetilde{q}^2 - \Lambda_1^2}\right)^{3/2}$$

Photon-valence quark coupling

$$F_i^D(q^2) = \frac{\Lambda_D^2}{\Lambda_D^2 - \widetilde{q}^2}\left(\frac{\Lambda_2^2}{\Lambda_2^2 - \widetilde{q}^2}\right)^i$$

QCD correction

$$\widetilde{q}^2 = q^2 \frac{\ln[(\Lambda_D^2 - q^2)/\Lambda_{\mathrm{QCD}}]}{\ln(\Lambda_D^2/\Lambda_{\mathrm{QCD}}^2)}$$

$\kappa_{V_{\mathrm{I}}}$ κ_{I}

- $\kappa_{V_{\mathrm{I}}}$ is the ratio of tensor to vector NNV_{I}-coupling at $q^2 = 0$
- Isospin anomalous magnetic moments: $\begin{cases} \kappa_{\mathrm{is}} = \kappa_p + \kappa_n \\ \kappa_{\mathrm{iv}} = \kappa_p - \kappa_n \end{cases}$

The ingredients...

$$F_i^{\mathrm{I}}(q^2) = \sum_{V_{\mathrm{I}}} \kappa_{V_{\mathrm{I}}}^{i-1} BW_{V_{\mathrm{I}}}(q^2) F_i^{V_{\mathrm{I}}}(q^2) + \left[\kappa_{\mathrm{I}}^{i-1} - \sum_{V_{\mathrm{I}}} \kappa_{V_{\mathrm{I}}}^{i-1} BW_{V_{\mathrm{I}}}(0) F_i^{V_{\mathrm{I}}}(0)\right] F_i^D(q^2)$$

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$V_I = \rho, \rho', \omega, \omega'$

$$F_i^{V_{\mathrm{I}}}(q^2) = f_i(q^2) = \left(\frac{\Lambda_1^2}{\Lambda_1^2 - \tilde{q}^2}\right)\left(\frac{\Lambda_2^2}{\Lambda_2^2 - \tilde{q}^2}\right)^i$$

$$F_1^{\phi}(q^2) = f_1(q^2)\left(\frac{\tilde{q}^2}{\tilde{q}^2 - \Lambda_1^2}\right)^{3/2}$$

$$F_2^{\phi}(q^2) = f_2(q^2)\left(\frac{\Lambda_1^2}{\mu_\phi^2}\frac{\tilde{q}^2 - \mu_\phi^2}{\tilde{q}^2 - \Lambda_1^2}\right)^{3/2}$$

Photon-valence quark coupling

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QCD correction

$$\tilde{q}^2 = q^2 \frac{\ln[(\Lambda_D^2 - q^2)/\Lambda_{\mathrm{QCD}}]}{\ln(\Lambda_D^2/\Lambda_{\mathrm{QCD}}^2)}$$

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The ingredients...

$$F_i^I(q^2) = \sum_{V_I} \kappa_{V_I}^{i-1} BW_{V_I}(q^2) F_i^{V_I}(q^2) + \left[\kappa_I^{i-1} - \sum_{V_I} \kappa_{V_I}^{i-1} BW_{V_I}(0) F_i^{V_I}(0)\right] F_i^D(q^2)$$

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$$BW_{V_I}(q^2) = \frac{g_{V_I} M_{V_I}^2}{f_{V_I}} \times [V_I\text{-propagator}] \begin{cases} (M_{V_I}^2 - q^2)^{-1} & V_I = \omega, \phi \\ [\text{analytic}] & V_I = \rho, \rho', \omega' \end{cases}$$

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$V_I = \rho,\ \rho',\ \omega,\ \omega'$

$$F_i^{V_I}(q^2) = f_i(q^2) = \left(\frac{\Lambda_1^2}{\Lambda_1^2 - \tilde{q}^2}\right)\left(\frac{\Lambda_2^2}{\Lambda_2^2 - \tilde{q}^2}\right)^i$$

$$F_1^{\phi}(q^2) = f_1(q^2)\left(\frac{\tilde{q}^2}{\tilde{q}^2 - \Lambda_1^2}\right)^{3/2}$$

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Photon-valence quark coupling

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$$\tilde{q}^2 = q^2 \frac{\ln[(\Lambda_D^2 - q^2)/\Lambda_{QCD}]}{\ln(\Lambda_D^2/\Lambda_{QCD}^2)}$$

κ_{V_I} κ_I

- κ_{V_I} is the ratio of tensor to vector NNV_I-coupling at $q^2 = 0$
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The ingredients. . .

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$$F_i^{V_{\mathsf{I}}}(q^2) = f_i(q^2) = \left(\frac{\Lambda_1^2}{\Lambda_1^2 - \tilde{q}^2}\right) \left(\frac{\Lambda_2^2}{\Lambda_2^2 - \tilde{q}^2}\right)^i$$

$$F_1^{\phi}(q^2) = f_1(q^2) \left(\frac{\tilde{q}^2}{\tilde{q}^2 - \Lambda_1^2}\right)^{3/2}$$

$$F_2^{\phi}(q^2) = f_2(q^2) \left(\frac{\Lambda_1^2}{\mu_\phi^2} \frac{\tilde{q}^2 - \mu_\phi^2}{\tilde{q}^2 - \Lambda_1^2}\right)^{3/2}$$

Photon-valence quark coupling

$$F_i^D(q^2) = \frac{\Lambda_D^2}{\Lambda_D^2 - \tilde{q}^2} \left(\frac{\Lambda_2^2}{\Lambda_2^2 - \tilde{q}^2}\right)^i$$

QCD correction

$$\tilde{q}^2 = q^2 \frac{\ln[(\Lambda_D^2 - q^2)/\Lambda_{\mathrm{QCD}}]}{\ln(\Lambda_D^2/\Lambda_{\mathrm{QCD}}^2)}$$

$\kappa_{V_{\mathsf{I}}}$ κ_{I}

- $\kappa_{V_{\mathsf{I}}}$ is the ratio of tensor to vector NNV_{I}-coupling at $q^2 = 0$
- Isospin anomalous magnetic moments: $\begin{cases} \kappa_{\mathrm{is}} = \kappa_p + \kappa_n \\ \kappa_{\mathrm{iv}} = \kappa_p - \kappa_n \end{cases}$

Time-like extension of the model

Analytic Breit-Wigner formulas

Relativistic Breit-Wigner formula for an unstable particle with mass $\boldsymbol{M}$ and constant width $\boldsymbol{\Gamma}$

$$BW(s) = \frac{1}{M^2 - s - i\Gamma M}$$

In the s-complex plane

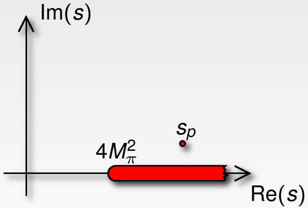
- Single pole
- No discontinuity cut

Energy-dependent widths are introduce considering decay rates extended to off shell particle masses. In the ρ meson case with $\boldsymbol{\Gamma(\rho \to \pi^+\pi^-) = \Gamma_\rho}$.

$$\Gamma(s) = \frac{\Gamma_s^0}{M_\rho} \frac{(s - 4M_\pi^2)^{3/2}}{s}$$

$$BW(s) = \frac{s}{s(M_\rho^2 - s) - i\Gamma_s^0(s - 4M_\pi^2)^{3/2}}$$

- Has the "required" discontinuity cut $(4M_\pi^2, \infty)$
- Maintains a complex pole $s_p \simeq M_\rho^2 + i\Gamma_\rho M_\rho$, slightly shifted w.r.t. the original position
- The power "3/2" in the denominator and the factor $1/s$ generate additional "physical" poles



Regularization of Breit-Wigner formulas

The BW formula with energy-dependent width has a set of N poles $\{z_j\}_{j=1}^{N}$ in the s-complex plane

- $P_N(s)$ is a suitable N degree polynomial
- β is a noninteger real number which defines the discontinuity cut

$$BW(s) = \frac{P_N(s)\prod_{j=1}^{N}(s-z_j)^{-1}}{M_\rho^2 - s - i\gamma(s-4M_\pi^2)^\beta}$$

To avoid unphysical divergences poles must be subtracted. BW formulas are **regularized** by adding counterparts that behave as the opposite of each pole.

Method #1

The subtraction can be done by hand...

$$\widetilde{BW}(s) = BW(s) - \sum_{k=1}^{N} \frac{P_N(z_k)\prod_{j=1,j\neq k}^{N}(z_k-z_j)^{-1}}{M_\rho^2 - z_k - i\gamma(z_k-4M_\pi^2)^\beta} \times \frac{1}{s-z_k}$$

- **Advantage**: easy to handle and to implement in codes
- **Drawback**: we need to know the pole positions

Regularization using dispersion relations

If $f(z)$ is an analytic function in the whole z complex plane with a real positive cut (s_0, ∞) and $f(z) = o(1/\ln|z|)$ as $z \to \infty$ then ....

$$f(z) = \frac{1}{\pi} \int_{s_0}^{\infty} \frac{\mathrm{Im}\,[f(x)]\,dx}{x - z}$$

If the function $f(z)$ has also a finite number of isolated poles $\{z_j\}_{j=1}^{N}$, **residues** have to be considered.....................

$$f(z) + 2\pi i \sum_{j=1}^{N} \mathrm{Res}\left[\frac{f(z')}{z' - z}, z_j\right] = \frac{1}{\pi} \int_{s_0}^{\infty} \frac{\mathrm{Im}\,[f(x)]\,dx}{x - z}$$

Writing

$$f(z) = \phi(z) \prod_{j=1}^{N} \frac{1}{z - z_j}$$

$\phi(z)$ is the pole-free part of $f(z)$

$$f(z) + \sum_{k=1}^{N} \frac{\phi(z_k)}{z_k - z} \prod_{j \neq k}^{N} \frac{1}{z_k - z_j} = \frac{1}{\pi} \int_{s_0}^{\infty} \frac{\mathrm{Im}\,[f(x)]\,dx}{x - z}$$

Method #2

$$\widetilde{BW}(s) = \frac{1}{\pi} \int_{4M_\pi^2}^{\infty} \frac{\mathrm{Im}\,[BW(s')]\,ds'}{s' - s}$$

- **Advantage**: we do not need to know the coordinates of the poles
- **Drawback**: we have to compute integrals over infinite intervals

Fit, results and dicussion

The fitting procedure

Space-like region

Quantity (Q_i)	G_M^p	G_E^p	G_M^n	G_E^n	$\frac{\mu_p G_E^p}{G_M^p}$	$\frac{\mu_n G_E^n}{G_M^n}$
n. of points (N_i)	68	36	65	14	25	13

Time-like

$\lvert G_{\text{eff}}^p \rvert$	$\lvert G_{\text{eff}}^n \rvert$
81(43)	5

Global χ^2

$$\chi^2 = \sum_{i=1}^{9} \tau_i \cdot \chi_i^2$$

Contribution of the data set: $\{q_k^2, v_k^i, \delta v_k^i; N_i\}$

$$\chi_i^2 = \sum_{k=1}^{N_i} \left(\frac{Q_i(q_k^2) - v_k^i}{\delta v_k^i} \right)^2$$

13 Free parameters

$\Lambda_1, \Lambda_2, \Lambda_D$

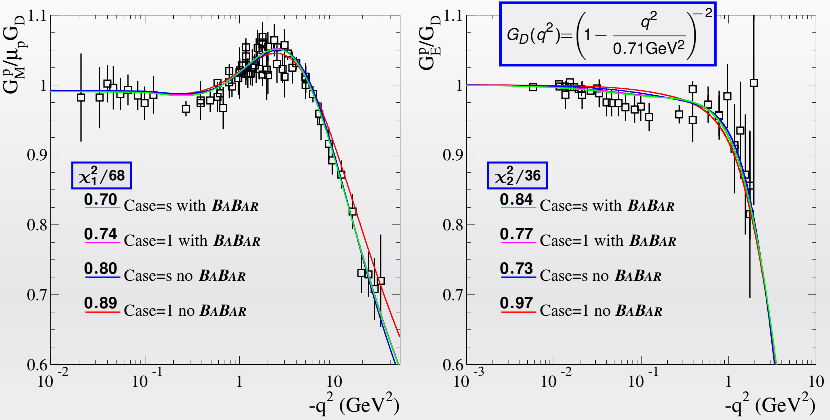
Parametrize hadronic FFs and control the transition from non perturbative to perturbative QCD

Five pairs $(\kappa_{V_l}, g_{V_l}/f_{V_l})$ of vector meson anomalous magnetic momenta and couplings with: $V_l = \rho, \rho', \omega, \omega', \phi$

- Masses and widths of all vector mesons are fixed to the PDG values
- For the QCD scale two values have been considered: $\Lambda_{QCD} = 0.15,\ 0.10$ GeV

Proton space-like G_E^p and G_M^p

Data collection: PRC82,045211



Two BW analytic formulas

- **Case=s**: $\Gamma_s(s) = \frac{\Gamma_s^0}{M}\frac{(s-s_{\rm th})^{3/2}}{s}$
- **Case=1**: $\Gamma_1(s) = \frac{\Gamma_1^0}{M^3}(s-s_{\rm th})^{3/2}$

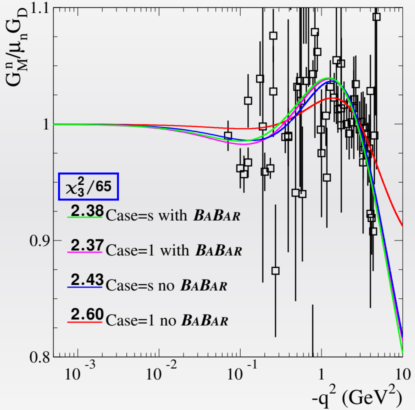
Two data sets: including and not ***BABAR*** time-like data on $|G_{\rm eff}^p(q^2)|$ extracted from the radiative process

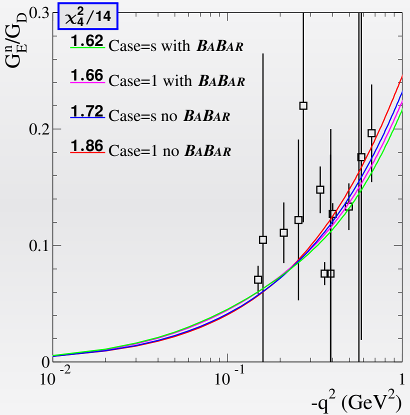
$$e^+e^- \to e^+e^-\gamma_{\rm init} \to p\bar{p}\gamma_{\rm init}$$

- with ***BABAR***
- no ***BABAR***

Neutron space-like G_E^n and G_M^n

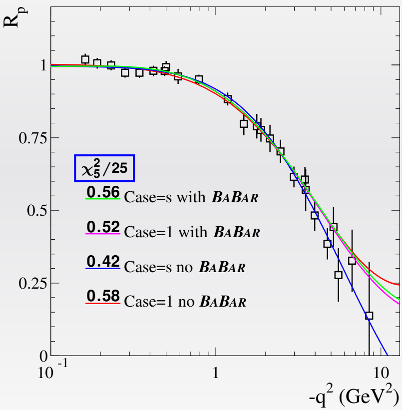
Data collection: PRC82,045211

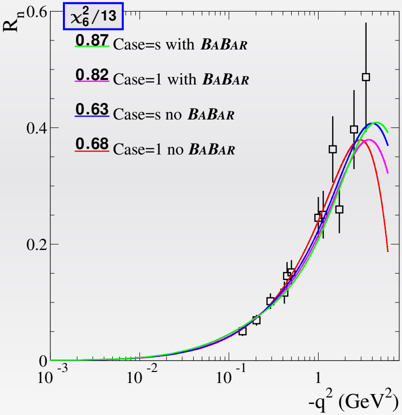




Space-like ratios
$R_p = \mu_p G_E^p / G_M^p$ and $R_n = \mu_n G_M^n / G_M^n$

Data collection: PRC82,045211





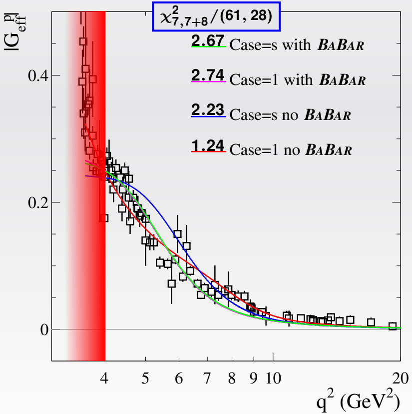
Time-like "effective" FFs: $|G^p_{\rm eff}|$ and $|G^n_{\rm eff}|$

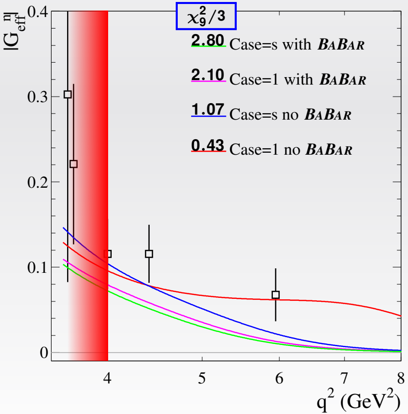
From experiments

$$|G^N_{\rm eff}| = \sqrt{\frac{\sigma(e^+e^- \to N\overline{N})}{\frac{4\pi\alpha^2}{3q^2}\sqrt{1-\frac{4M_N^2}{q^2}}\left(1+\frac{2M_N^2}{q^2}\right)}}$$

Fitting function

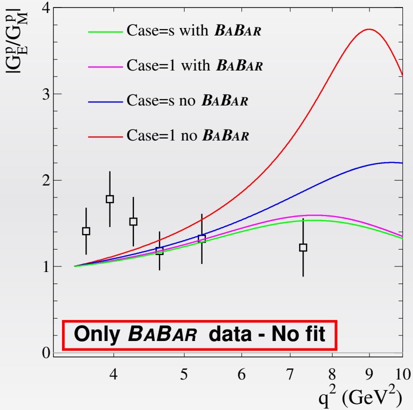
$$|G^N_{\rm eff}| = \sqrt{\frac{|G^N_M(q^2)|^2 + \frac{2M_N^2}{q^2}|G^N_E(q^2)|^2}{1+\frac{2M_N^2}{q^2}}}$$

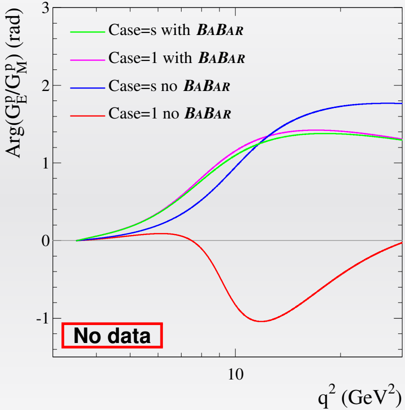




Prediction for the time-like ratio G_E^p / G_M^p

Data: PRD73,012005





Summary of χ^2 contributions

	Q_i	N_i	minimum χ^2_i			
			case=s With BaBar	case=1 With BaBar	case=s No BaBar	case=1 No BaBar
space-like	G_M^p	68	48.7	50.1	54.6	60.8
	G_E^p	36	30.4	27.6	26.2	35.0
	G_M^n	65	154.6	154.2	158.2	167.0
	G_E^n	14	22.7	23.2	24.1	26.0
	$\mu_p G_E^p / G_M^p$	25	13.9	12.9	10.6	14.4
	$\mu_n G_E^n / G_M^n$	13	11.3	10.7	8.2	8.9
time-like	$\lvert G_{\rm eff}^p \rvert$	61 (28)	162.5	166.7	62.2	35.0
	$\lvert G_{\rm eff}^n \rvert$	3	8.4	6.3	3.2	0.3
	Total	285(252)	452.5	451.7	347.3	347.4
	Normalized χ^2		1.663	1.661	1.453	1.454

Best values of fit parameters and constants

Parameter	case = s With *BaBar*	case = 1 With *BaBar*	case = s No *BaBar*	case = 1 No *BaBar*
g_ρ/f_ρ	2.766	2.410	0.9029	0.4181
κ_ρ	-1.194	-1.084	0.8267	0.6885
M_ρ (GeV)	0.7755 (fixed)			
Γ_ρ (GeV)	0.1491 (fixed)			
g_ω/f_ω	-1.057	-1.043	-0.2308	-0.4894
κ_ω	-3.240	-3.317	-9.859	-1.398
M_ω (GeV)	0.78263 (fixed)			
g_ϕ/f_ϕ	0.1871	0.1445	-0.0131	-0.1156
κ_ϕ	-2.004	-3.045	37.218	-0.2613
M_ϕ (GeV)	1.019 (fixed)			
μ_ϕ (GeV)	20.0 (fixed)			
$g_{\omega'}/f_{\omega'}$	2.015	1.974	1.265	1.649
$\kappa_{\omega'}$	-2.053	-2.010	-2.044	-0.6712
$M_{\omega'}$ (GeV)	1.425 (fixed)			
$\Gamma_{\omega'}(GeV)$	0.215 (fixed)			
$g_{\rho'}/f_{\rho'}$	-3.475	-3.274	-0.8730	-0.0369
$\kappa_{\rho'}$	-1.657	-1.724	-2.832	-104.35
$M_{\rho'}$ (GeV)	1.465 (fixed)			
$\Gamma_{\rho'}$ (GeV)	0.400 (fixed)			
Λ_1 (GeV)	0.4801	0.5000	0.6474	0.6446
Λ_2 (GeV)	3.0536	3.0562	3.0872	3.6719
Λ_D (GeV)	0.7263	0.7416	0.8573	0.8967
Λ_{QCD} (GeV)	0.150			0.100

Discussion

Time-like extension of the Lomon-Gari-Krümpelman model

- Breit-Wigner formulas describing broad intermediate vector mesons have been modified including energy-dependent widths in two scenarios: $\Gamma_1(q^2)$, minimal alteration, and $\Gamma_s(q^2)$ derived from relativistic perturbation theory
- A **regularization procedure** has been defined to remove unwanted poles and so to fulfill the analyticity requirements in the whole q^2 complex plane

Fit results

- The improvements in the BW formulas do not affect the space-like fit quality
- The **simultaneous space-like and time-like** fit is satisfactory
- The χ^2 contributions from each space-like data set are almost **unchanged** between **case=1** and **case=*s***
- The quality of the fit is **poorer when *BABAR* data are included**

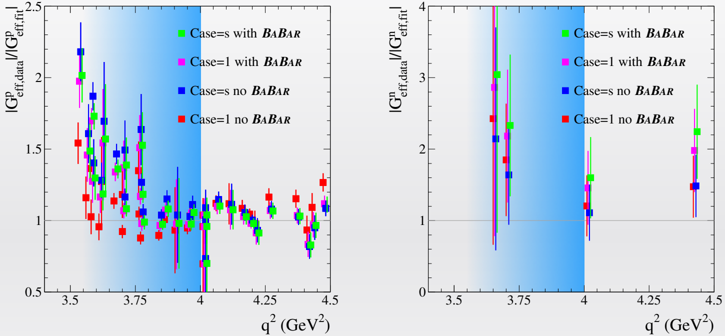
Predictions

- The knowledge of the analytic structure of EMFFs allows us to make predictions on **polarization observables** like EMFF **phases** and **"pure" moduli**
- Measurements of such observables would be effective in **discriminating among the different models and parametrizations**

BACK-UP SLIDES

Time-like threshold behavior

$$|G_{\rm eff}^N| = \sqrt{\sigma(e^+e^- \to N\overline{N})} \left[\frac{4\pi\alpha^2}{3q^2}\sqrt{1-4M_N^2/q^2}\left(1+2M_N^2/q^2\right)\right]^{-1/2}$$



- As a consequence of near-threshold **flat cross sections**, effective nucleon EMFFs have a steep enhancement when $q^2 \to (2M_N)^2$
- Such a flat cross section is in contrast with the expectation in case of smooth EMFFs ($\sigma \propto \beta_N$)
- To extract Born cross section and hence $|G_{\rm eff}^N|$ in the threshold region, data have to be corrected for **Coulomb** as well as **strong** effects

To avoid ambiguities due to the not well known form and interplay of these threshold corrections, **time-like data below $q^2 = 4$ GeV^2 have not been considered**