



XX Baldin ISHEPP
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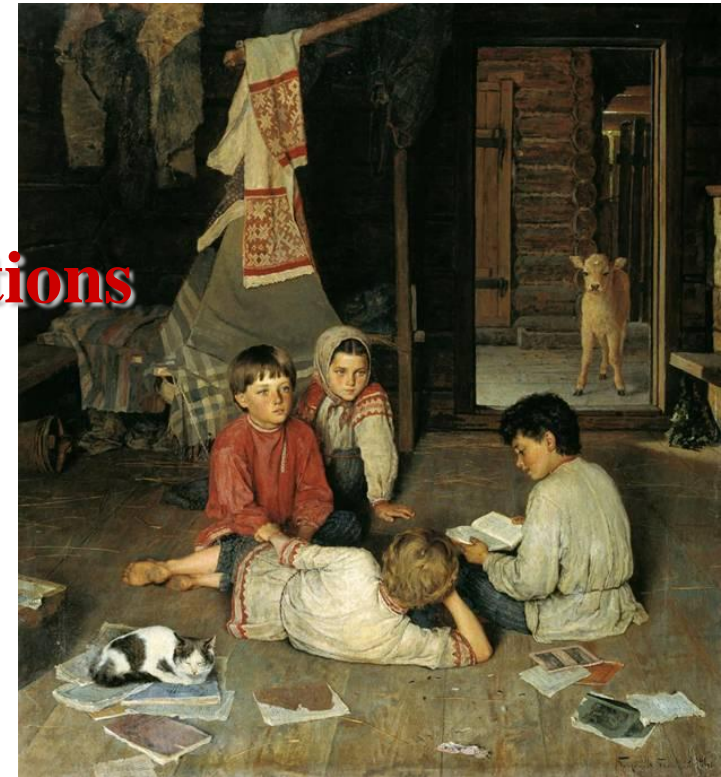
**Modification of fundamental interactions
in strong electromagnetic fields**

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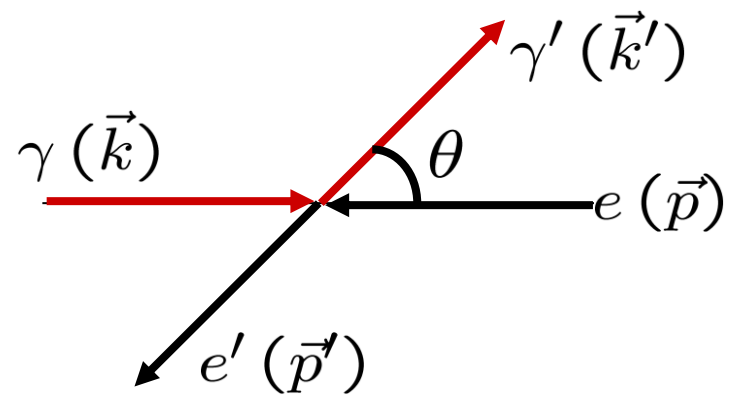


Bogdanov-Belsky
“New tale”

Outline

- Compton $\gamma e \rightarrow \gamma e$ scattering (Klein-Nishina equation)
- Volkov solution of Dirac equation in strong EM field
- Emission of photons by an electron in a strong EM field
- Reaction $e \rightarrow e + 2$ neutrinos in a strong EM field
- Weak decay of a neutron in a strong EM field
- Summary

Compton $\gamma e \rightarrow \gamma e$ Scattering (Klein-Nishina equation)



$$d\sigma = \frac{1}{16\pi(s - M_e^2)^2} |T|^2 dt$$

with $s = (p + k)^2 = M_e^2 + 2E_\gamma(E + p)$
 $t = (p - p')^2 = (k' - k)^2$

$$T_{\gamma e \rightarrow \gamma e} \rightarrow \text{[Feynman diagrams]} + \text{[Feynman diagrams]} = e^2 \epsilon_\mu^*(\gamma') \epsilon_\nu(\gamma) \cdot [\bar{u}(e') M^{\mu\nu} u(p)] \times (2\pi)^4 \delta(p + k - p' - k')$$

$$M^{\mu\nu} = \gamma^\mu \frac{\gamma \cdot p + \gamma \cdot k + M_e}{2p \cdot k} \gamma^\nu + \gamma^\nu \frac{\gamma \cdot k - \gamma \cdot p' + M_e}{2p' \cdot k} \gamma^\mu$$

$$d\sigma = \frac{1}{16\pi(s - M_e^2)^2} |T|^2 dt$$

$$|T|^2 = \frac{e^2}{4} \sum_{\lambda_\gamma, \lambda'_\gamma, m_e, m'_e} |\epsilon_\mu^*(\gamma') [\bar{u}(e') M^{\mu\nu} u(p)] \epsilon_\nu(\gamma)|^2$$

$$d\sigma/dt = \pi r_0^2 F(p, p', k) M_e^2 / (s - M_e^2)^2,$$

with

$$r_0 = \frac{e^2}{4\pi M_e} \equiv \frac{\alpha}{M_e} \simeq 2.82 \text{ fm} \quad \text{“electron radius”}$$

and function

$$F(p, p', k) = \left\{ \left(\frac{M_e^2}{2k \cdot p} + \frac{M_e^2}{2k \cdot p'} \right)^2 + \left(\frac{M_e^2}{2k \cdot p} + \frac{M_e^2}{2k \cdot p'} \right) - \frac{1}{4} \left(\frac{k \cdot p}{k \cdot p'} + \frac{k \cdot p'}{k \cdot p} \right) \right\}$$

relation between cross section $d\sigma$ to probability dW

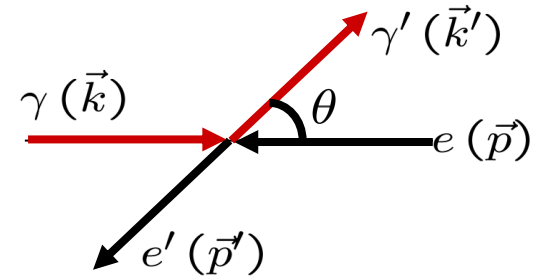
$$dW = J \cdot \frac{1}{V_\gamma} \cdot d\sigma$$

probability

flux factor

photon density ?

$$J = v(\gamma, e) = \frac{\sqrt{s}}{E_e}$$



$$\vec{A}_\gamma = \vec{a}_x \cos(\omega t - kz) + \vec{a}_y \sin(\omega t - kz)$$

circularly polarized photon field

$$\mathcal{E} = (E^2 + H^2)/2 = E^2 \quad \text{energy density of EM field}$$

$$\vec{E} = -\frac{\partial \vec{A}_\gamma}{\partial t} = \omega(\vec{a}_x \sin(\omega t - kz) - \vec{a}_y \cos(\omega t - kz))$$

$$\langle \vec{E}^2 \rangle = \omega^2 a^2; \quad \frac{1}{V_\gamma} = \frac{\mathcal{E}}{\omega} = \omega \cdot a^2 = \frac{\omega M_e^2 \xi^2}{4\pi\alpha}$$

$$a^2 = a_x^2 = a_y^2$$

with $\xi^2 = \frac{e^2 a^2}{M_e^2}$

reduced intensity of EM field

$$dW = d\sigma \cdot \frac{\sqrt{s}}{E_e} \cdot \frac{M_e^2 \xi^2 \omega_\gamma}{4\pi\alpha}$$

Final expression for the Compton scattering probability dW

$$dW = d\sigma \cdot \frac{\sqrt{s}}{E_e} \cdot \frac{M_e^2 \xi^2 \omega \gamma}{4\pi\alpha}$$

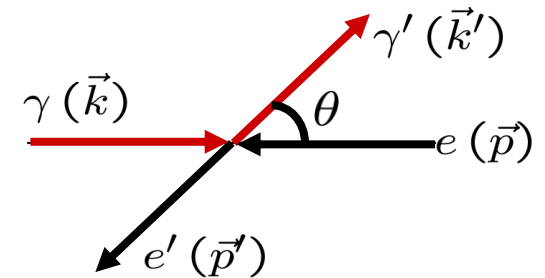
square of the total energy $s = (p + k)^2$

square of the momentum transfer

in cross channel $\bar{s} = (p - k')^2$

square of momentum transfer

$$t = (k - k')^2 = -2k \cdot k' = -2\omega\omega'(1 - \cos\theta) \quad (\text{or } \theta)$$

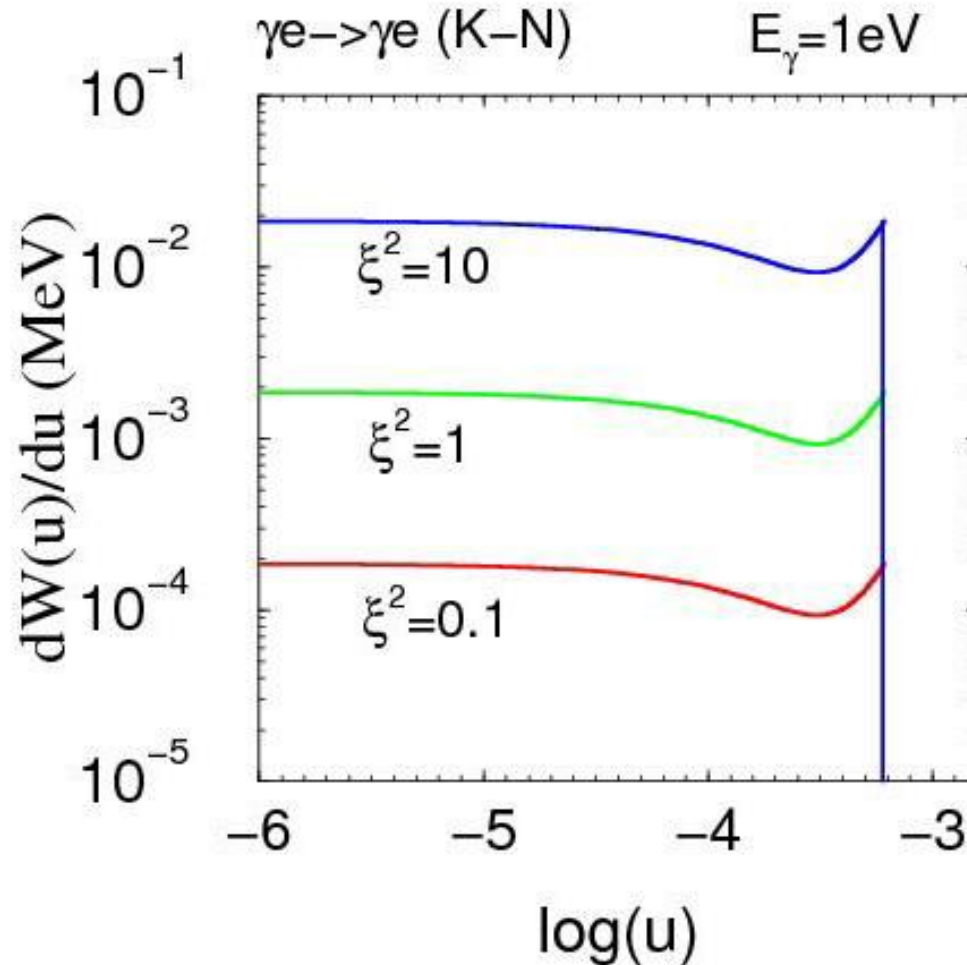


For further application it is convenient to use *new invariant variable u*

$$u = \frac{k \cdot k'}{k \cdot p'} = \frac{\omega_c(1 - \cos\theta_c)}{E'_e + \omega_c \cos\theta_c} \quad 0 \leq u \leq u_{\max} = \frac{2\omega_L}{M_e} (\ll 1)$$

$$dW = \pi r_0^2 F(p, p', k) \cdot \frac{M_e^2}{2E_e} \cdot \frac{M_e^2 \xi^2}{4\pi\alpha} \cdot \frac{du}{(1 + u)^2}$$

Compton $\gamma e \rightarrow \gamma e$ scattering (results)



Shape of differential distribution does not depend on field intensity because ξ^2 describes an overall factor BUT not matrix element!

Electron in a strong electromagnetic fields

D.M. Volkov, Z. Phys. **94**, 250 (1935)

Über eine Klasse von Lösungen der Diracschen Gleichung.

1. Der Fall eines sinusoidalen Feldes. — 2. Lösung für den Fall, daß das äußere Feld aus polarisierten, in einer bestimmten Richtung fortschreitenden Wellen besteht, die ein abzählbares Spektrum nach Frequenz und Anfangsphasen haben.

Dirac second order equation

$$[(\hat{p} - eA)^2 - m^2 - i\frac{1}{2}F_{\mu\nu}\sigma^{\mu\nu}]\psi = 0,$$

where $F_{\mu\nu} = \partial_\nu A_\mu - \partial_\mu A_\nu$ is EM field tensor

4-component spinor

$$\sigma_{\mu\nu} = \frac{i}{2}(\gamma_\mu\gamma_\nu - \gamma_\nu\gamma_\mu), \quad \gamma_\mu - 4 \times 4 \text{ Dirac matrices} \quad \gamma_0 = \begin{pmatrix} I & 0 \\ 0 & -I \end{pmatrix}, \quad \vec{\gamma} = \begin{pmatrix} 0 & \vec{\sigma} \\ -\vec{\sigma} & 0 \end{pmatrix}$$

and $A = (0, \vec{A}_\gamma)$ is four vector of electromagnetic field with the special part chosen as

$$\vec{A}_\gamma = \vec{a}_x \cos(\phi) + \vec{a}_y \sin(\phi); \quad \phi = k \cdot x = \omega t - kz$$

with $|\vec{a}_x| = |\vec{a}_y| = a$

Solution:

$$\psi_p = \left[1 + \frac{e(\gamma \cdot k)(\gamma \cdot A)}{2(k \cdot p)} \right] e^{iS'(\phi)} \frac{u_p}{\sqrt{2E_p}} e^{-ip \cdot x}$$

with

$$S'(\phi) = - \int_0^{kx} \left[\frac{e(p \cdot A)}{(k \cdot p)} - \frac{e^2 A^2}{2(k \cdot p)} \right] d\phi'$$

when $\vec{A} \rightarrow 0$ *or* $(a_x, a_y \rightarrow 0)$

$$\psi_p \rightarrow \frac{u_p}{\sqrt{2E_p}} e^{-ip \cdot x}$$

Dirac solution for free electron

Properties of Volkov's solution

Effective “quasi” momentum

$$\begin{aligned}\langle \psi^* (\hat{p}^\mu - eA^\mu) \psi \rangle &= q^\mu \\ &\equiv p^\mu - \frac{e^2 \bar{A}^2}{2(k \cdot p)} k^\mu = p^\mu + \frac{e^2 a^2}{2(k \cdot p)} k^\mu = p^\mu + \frac{\xi^2 m_e^2}{2(k \cdot p)} k^\mu \\ \bar{A}^2 &= -\frac{1}{2}(a_x^2 + a_y^2) = -a^2\end{aligned}$$

Effective electron mass

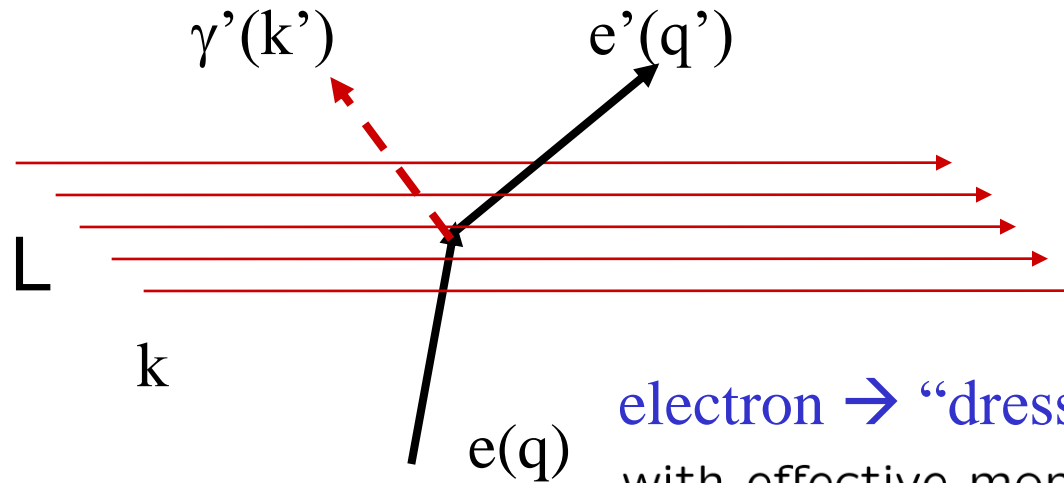
$$q^2 = m_*^2 \equiv m_e^2 \left(1 - \frac{e^2 \bar{A}^2}{m_e^2} \right) = m_e^2 (1 + \xi^2)$$

$$m_{e*}^2 = m^2 (1 + \xi^2)$$

$$\text{with } \xi^2 = \frac{e^2 a^2}{m_e^2} = \frac{e^2 E^2}{m_e^2 \omega^2}$$

“quasi-momentum” and effective mass define momentum-energy conservation

Emission of a photon by an electron in the field of a strong electromagnetic wave



electron \rightarrow “dressed” electron
with effective momentum $p \rightarrow q$
and mass $m_e^2 \rightarrow m_e^2(1 + \xi^2)$

$$T_{fi} = -ie \int \psi_f^* (\gamma \cdot \varepsilon_f^*) \psi_i e^{ik'x} \frac{d^4x}{\sqrt{2\omega'}}$$

Interaction of electron with an external field is considered non-perturbatively

Interaction of electron with outgoing photon is considered in first order of perturbation theory

Structure of matrix element

$$T_{fi} = -ie \int \psi_f^*(\gamma \cdot \varepsilon_f^*) \psi_i e^{ik'x} \frac{d^4x}{\sqrt{2\omega'}}$$

$$\frac{\bar{u}_{p'}}{\sqrt{2E_{p'}}} e^{-ip' \cdot x} \left[1 + \frac{e(\gamma \cdot A)(\gamma \cdot k)}{2(k \cdot p')} \right] e^{iS'(k \cdot x, p'_e)}$$

$$\left[1 + \frac{e(\gamma \cdot k)(\gamma \cdot A)}{2(k \cdot p)} \right] e^{iS'(k \cdot x, p_e)} \frac{u_p}{\sqrt{2E_p}} e^{-ip \cdot x}$$

“non-perturbative” outgoing
electron

“non-perturbative” incoming
electron

$$\rightarrow \frac{-ie}{\sqrt{2q_0 2q'_0 2\omega'}} \int M(kx) e^{-i(q - q' - k')x} d^4x$$

with

$$M(kx) = [..]_f \bar{u}_{p'}(\gamma \cdot \varepsilon_f^*) [..]_i u_p e^{-i(S(kx) - S'(kx))}$$

In “regular” Compton scattering $\gamma e \rightarrow \gamma e$,
one has

$$T \sim \int M(k, k', p, p') e^{-i(p+k-p'-k')x} d^4x \neq (2\pi)^4 \delta^4(q+k-q'-k') \cdot M$$

$$= (2\pi)^4 \delta^4(p+k-p'-k') M(k, k', p, p')$$

Structure of matrix element (continuing)

$$\frac{-ie}{\sqrt{2q_0 2q'_0 2\omega'}} \int M(kx) e^{-i(q-q'-k')x} d^4x$$

Fourier series

$$M(kx) = \sum_{n=-\infty}^{\infty} e^{-in kx} M_n(k, k', q, q')$$

The amplitude becomes a sum of infinite numbers of partial harmonics

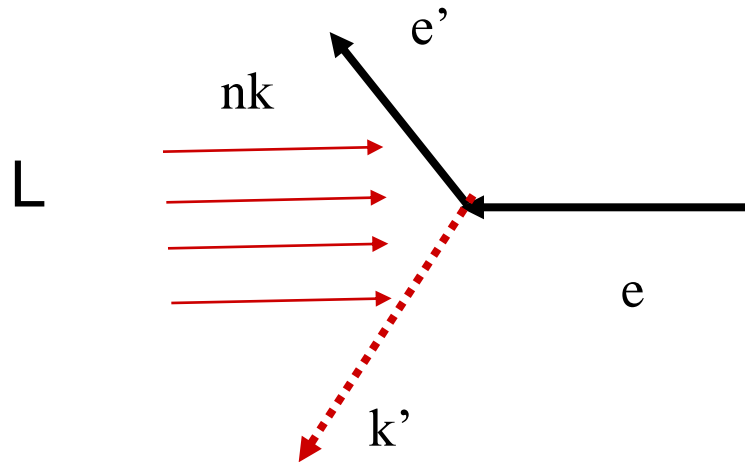
$$\begin{aligned} T_{fi} &= -ie \sum_{n=-\infty}^{\infty} M_n \int e^{-i(q+nk-q'-k')} d^4x \\ &= \sum_n -ie M_n (2\pi)^4 \delta^4(q + nk - q' - k') \end{aligned}$$

Each harmonic describes absorption (emitting) of n photons of external field \mathbf{A} with wave vector k and emitting of outgoing photon with the wave vector k' with corresponding conservation law

Probability is a sum of partial contributions

$$dW = \sum_n dW_n$$

$$dW_n = \frac{1}{16\pi E_q} |T_n|^2 \frac{du}{(1+u)^2}$$

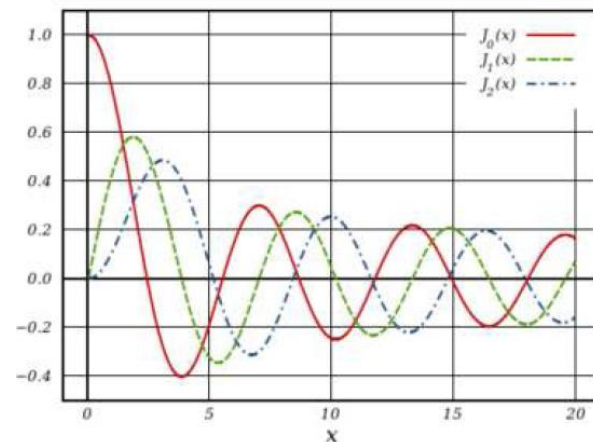


Properties of partial contributions

$$dW_n = \frac{\alpha}{4E_q} \frac{du}{(1+u)^2} \left\{ -4J_n^2(z) + \xi^2 \left(2 + \frac{u^2}{1+u} \right) (J_{n+1}^2(z) + J_{n-1}^2(z) - 2J_n^2(z)) \right\}$$

$$J_n(z) = \int_{-\pi}^{\pi} e^{i(n\phi - z \sin \phi)} d\phi; \quad \phi = kx$$

$$u_{\max} = \frac{2n\omega_L(E_e + p)}{m_e^2(1 + \xi^2)} \simeq \begin{cases} \frac{n\omega_\gamma^L}{m_e\sqrt{1+\xi^2}} & \text{for } E_e \simeq m_e \\ \frac{4nE_e\omega_\gamma^L}{m_e^2(1+\xi^2)} & \text{for } E_e \gg m_e \end{cases}$$

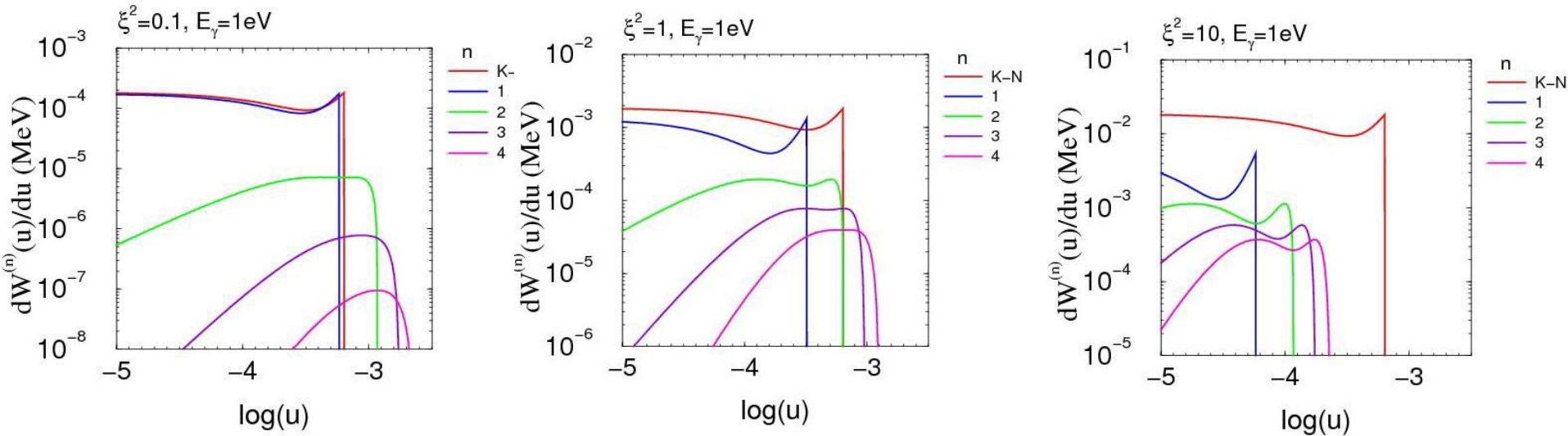


Kinematical limit (phase space) **increases** ($n > 0$)

2 effects: { Electron can interact with a few photons simultaneously (“cumulative” effect)
Dressed electron mass exceeds free electron mass

Results in **decrease** of the phase space even for one photon absorption

Photon emission in strong EM (results)



At small field intensity $\xi^2 \ll 1$ effect of mass modification is small, “cumulative” effect is large

At large field intensity $\xi^2 \gg 1$ effect of mass modification is larger, than “cumulative” effect. However, the later one is also important.

At $\xi^2 \geq 1$ standard Klein-Nishina equation does not work even for $n=1$.

pioneering works:

N.~D.~Sengupta, Bull. Calcutta Math. Soc. {\bf 44}, 175, (1952).

I.~I.~Goldman, Phys. Lett. {\bf 8}, 103 (1964)

L.~S.~Brown and T. W. B. Klibbe, Phys. Rev. A {\bf 133}, 705 (1964).

H.~R.~Reiss, J. Math. Phys. {\bf 3}, 59 (1962).

A.~I.~Nikishov and V.~I.~Ritus, Sov. Phys. JETP {(1964-79)}

N.~B.~Narozhnyi, A.~I.~Nikishov, and V.~I.~Ritus,
Sov. Phys. JETP {\bf 20}, 622 (1965)
[Zh. Eksp. Teor. Fiz. {\bf 47}, 930 (1964)].

V.~A.~Lyulka, Zh. Eksp. Teor. Fiz. {\bf 69}, 800 (1975).

N.~P.~Merenkov, Yad. Fiz. {\bf 42}, 1484 (1985).

V.~V.~Skobelev, Yad. Fiz. {\bf 46}, 1738 (1987).

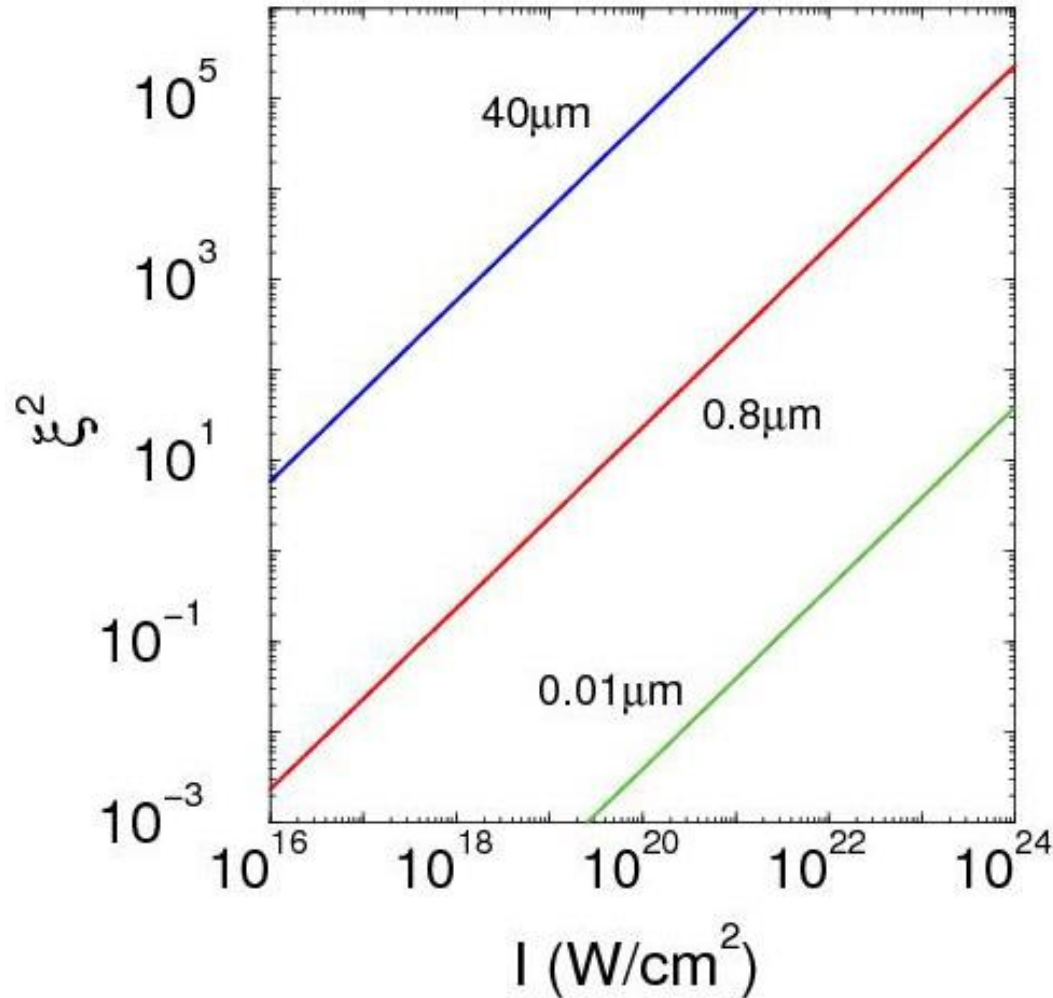
$$\xi^2 \ll 1$$

$$\xi^2 \rightarrow \infty;$$

$$\xi \sim \frac{1}{\omega\gamma}$$

$$\chi = \xi \frac{kp}{m_e^2}$$

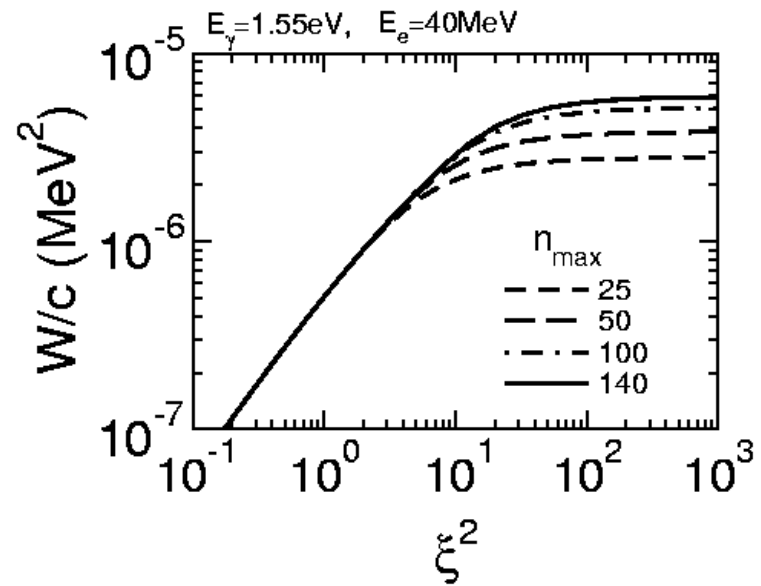
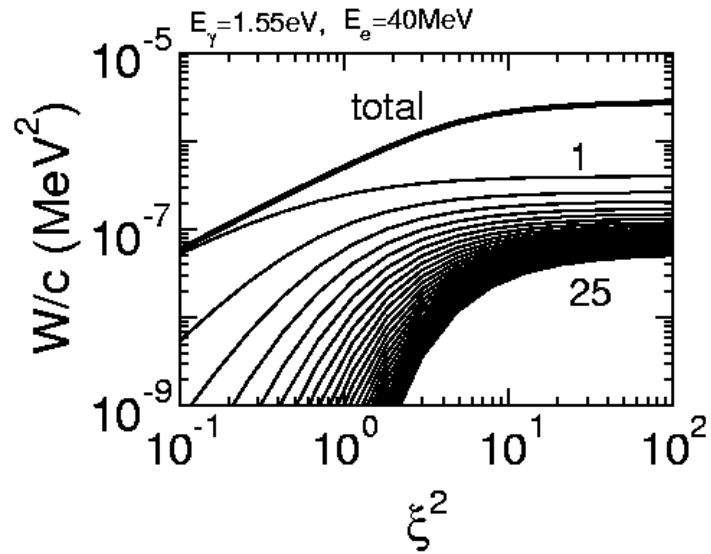
Dependance of reduced field strength ξ^2 on laser pulse intensity I at different wavelength λ



$$\xi^2 = \frac{4\pi\alpha(\hbar c)^2}{m_e^2 \omega_\gamma^2 c} I$$

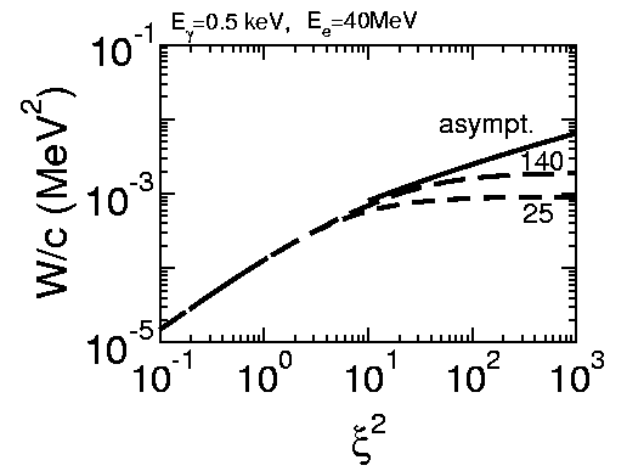
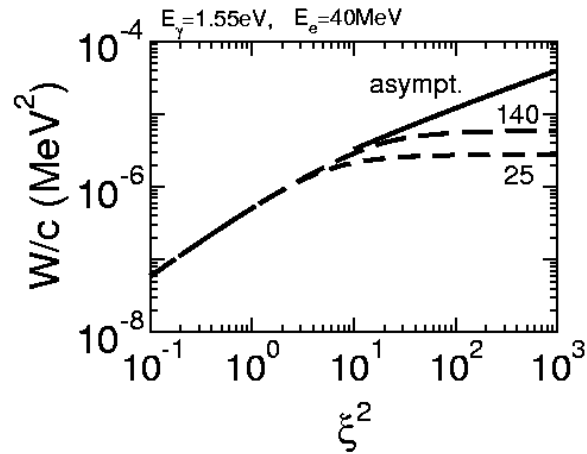
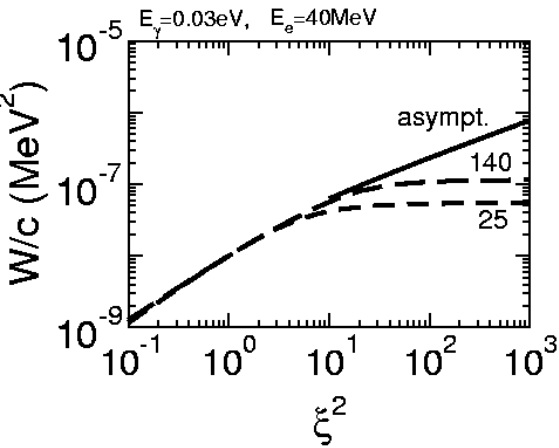
$$\lambda = \frac{2\pi\hbar c}{\omega_\gamma}$$

Problem of convergence



Asymptotic method

$$\sum_n \rightarrow \int_{n_{\min}(\xi)}^{\infty} dn$$



Summary of this part:

Non-perturbative effects of QED may be seen even at small $\xi^2 \simeq 0.1$

Perturbative QED does not work at finite values of $\xi^2 \geq 1$

Difference between predictions of perturbative QED and non-perturbative QED is large both qualitatively and quantitatively

Emission of neutrino pairs by an electron in a field of strong electromagnetic wave

Neutrino emission in $e \rightarrow e' + \bar{\nu}\nu$ is forbidden by energy arguments

$$P_i^2 = P_f^2 \rightarrow p_e = p_{e'} + p_{2\nu}$$

$$M_e^2 = M_{e'}^2 + M_{2\nu}^2 + 2p_{e'} \cdot p_{2\nu}$$

$$E_{e'}E_{2\nu} + M_{2\nu}^2/2 = |\vec{p}_{e'}| \cdot |\vec{p}_{2\nu}| \cos \theta$$

It is wrong, because

$$E_{e'} > |\vec{p}_{e'}|, \quad E_{2\nu} > |\vec{p}_{2\nu}|$$

and therefore, always

$$E_{e'}E_{2\nu} + M_{2\nu}^2/2 > |\vec{p}_{e'}| \cdot |\vec{p}_{2\nu}| \cos \theta$$

Reaction $\gamma e \rightarrow e' + \bar{\nu}\nu$ **is allowed. It is an analog to the Compton scattering**

Three types of neutrino may be emitted:

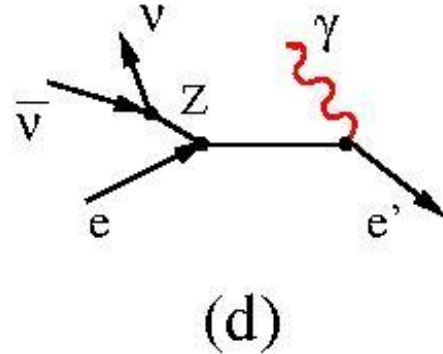
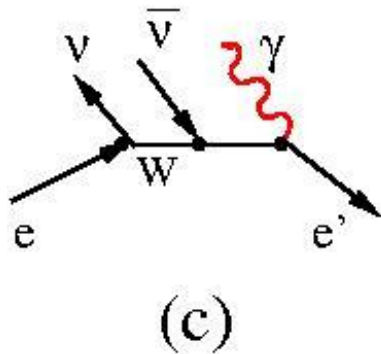
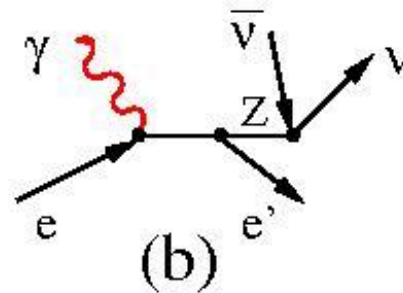
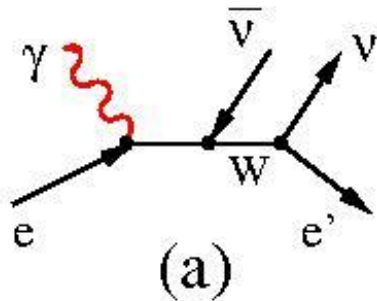
$$\gamma e \rightarrow e' + \bar{\nu}_e \nu_e \text{ (W, Z)}$$

$$\gamma e \rightarrow e' + \bar{\nu}_\mu \nu_\mu \text{ (Z)}$$

$$\gamma e \rightarrow e' + \bar{\nu}_\tau \nu_\tau \text{ (Z)}$$

Only

$\bar{\nu}_e \nu_e$



All three types

$\bar{\nu}_e \nu_e$

$\bar{\nu}_\mu \nu_\mu$

$\bar{\nu}_\tau \nu_\tau$

Question: how is neutrino emission modified in strong EM field?

electron modification is important

$$\begin{cases} \psi^D \rightarrow \psi^V \\ p \rightarrow q \\ m_e^2 \rightarrow m_*^2 = m_e^2 + e^2 a^2 = m_e^2 + m_e^2 \xi^2 \end{cases}$$

W^\pm modification is negligible:

$$\begin{cases} m_W^2 \rightarrow m_{W*}^2 = m_W^2 + e^2 a^2 = m_W^2 + m_e^2 \xi^2 \\ \Delta m_W = \frac{\Delta m_W^2}{2m_W} = m_e \cdot \xi^2 \frac{m_e}{m_W} \sim m_e \times 10^{-5} \end{cases}$$

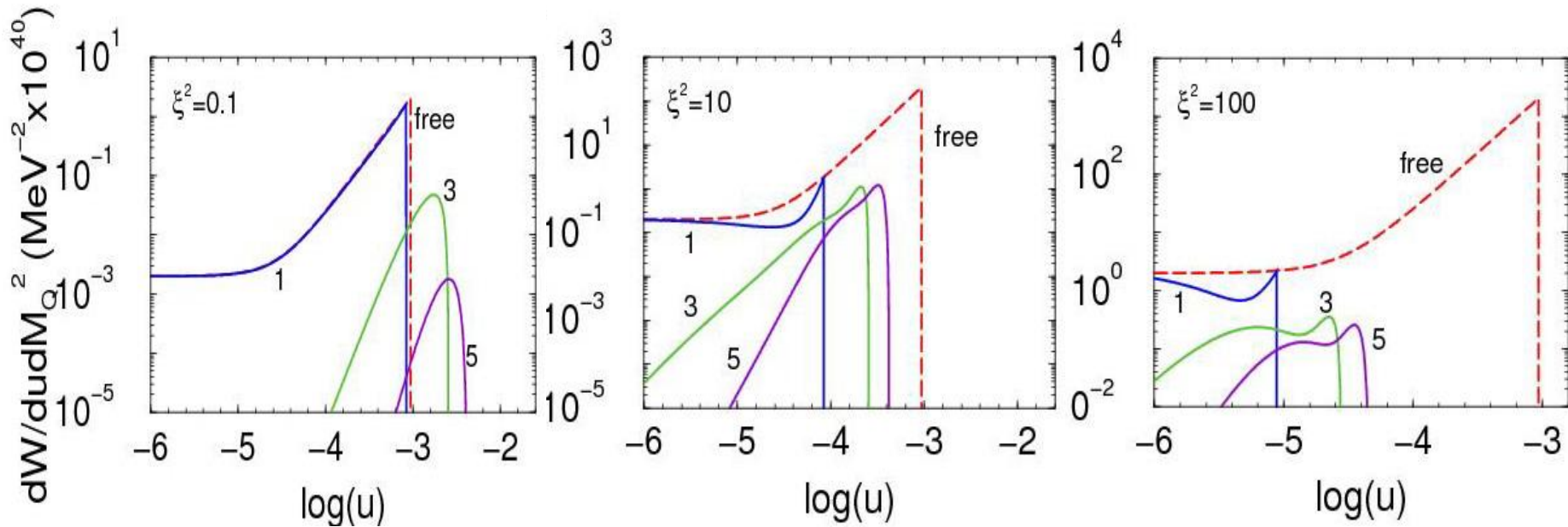
Amplitude(s)

$$T_{fi} = \frac{G_F}{\sqrt{2}} \int [\psi_f^* \gamma_\mu (a - b\gamma_5) \psi_i] \otimes [u_\nu \gamma^\mu (1 - \gamma_5) v_{\bar{\nu}}] e^{i(k_\nu + k_{\bar{\nu}})x} \frac{d^4x}{\sqrt{2E_\nu 2E_{\bar{\nu}}}}$$

Probability

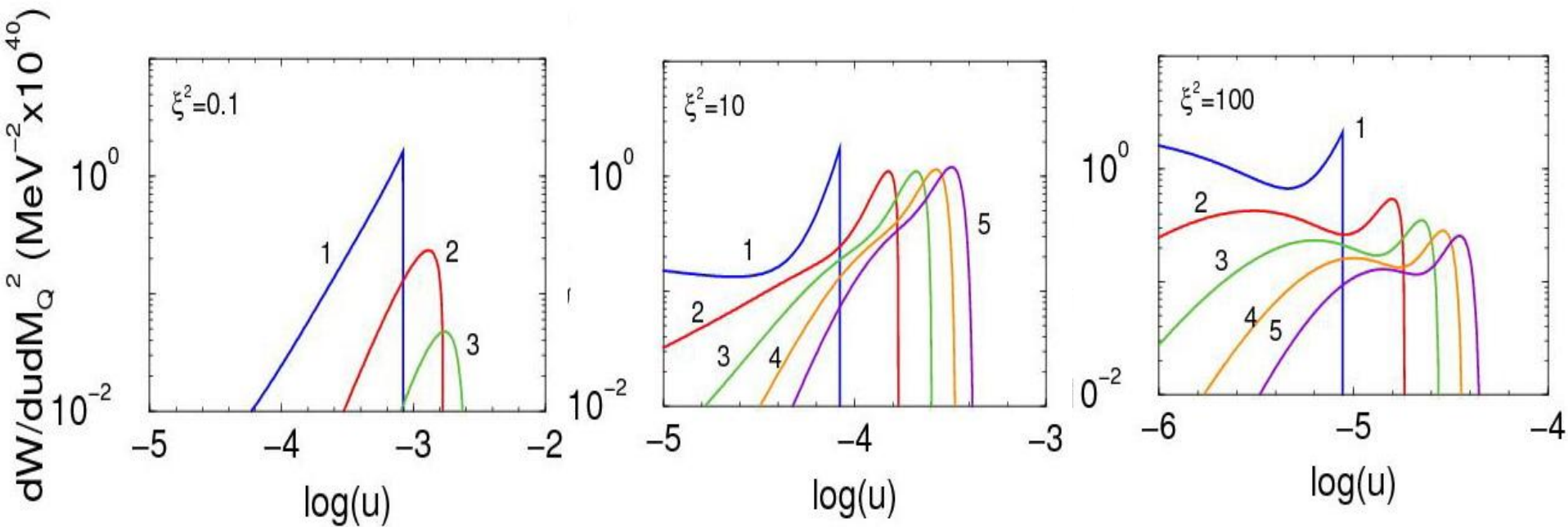
$$dW = \sum_{n>0} dW_n$$

2 neutrinos emission in strong EM (results)



- ✦ *Variation of ξ^2 in $\gamma e \rightarrow e \nu_e \bar{\nu}_e$ reaction in vacuum does not change shape of distribution. It changes overall normalization*
- ✦ *Increase of ξ^2 leads to decrease of kinematical limit (see $n=1$)*
- ✦ *“Cumulative effect” – reactions with $n>1$ increases the phase space and the kinematical limit*

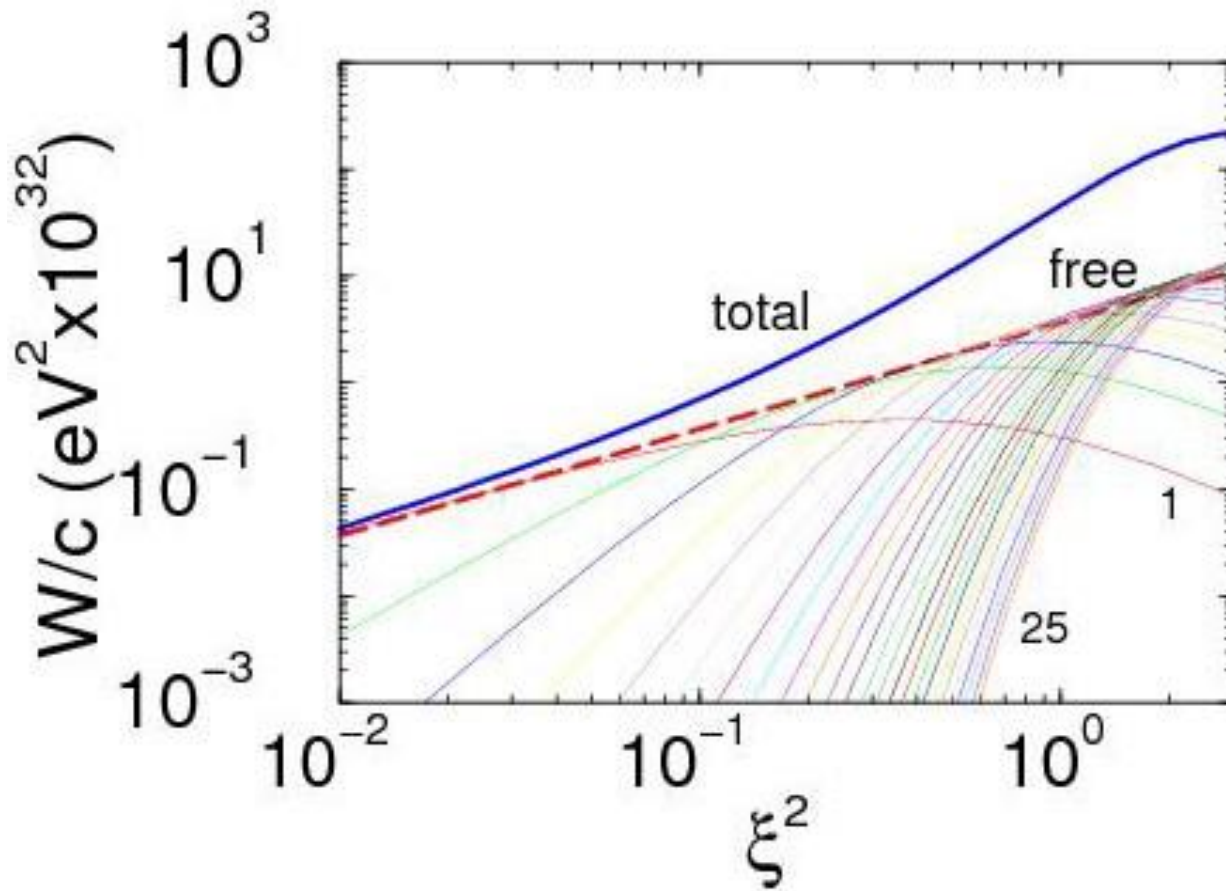
2 neutrinos emission in strong EM (results, without $\gamma e \rightarrow 2 \nu$, continuing)



In general, higher harmonics are not suppressed at large ξ^2

Total yield of neutrino as a function of ξ^2

$$\omega_\gamma = 1.56 \text{ eV}, E_e = 40 \text{ MeV}$$



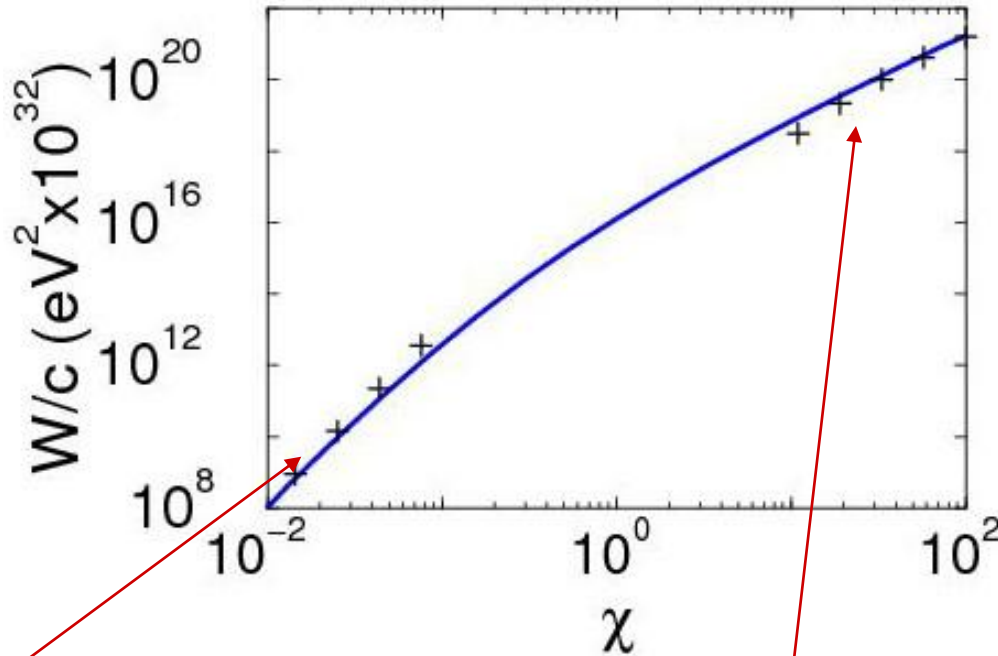
Higher harmonics dominate at large ξ^2 : $n_{\text{max}} = f(\xi^2)$

$$W(\chi, \xi^2)$$

$$\chi = \frac{e\sqrt{(\langle F_{\mu\nu} p_e^\nu \rangle)^2}}{m_e^3} \geq 0, \quad \xi^2 = -\frac{e^2 \langle A \rangle^2}{m_e^2} \geq 0$$

$$W(\chi, \xi^2 \rightarrow \infty) \rightarrow W_A(\chi)$$

Asymptotic total yield of neutrino in limit of $\xi^2 \rightarrow \infty$



$$\sim \frac{G_F^2 m_e^6}{192\sqrt{3}\pi^3} \chi^5, \quad \chi \ll 1$$

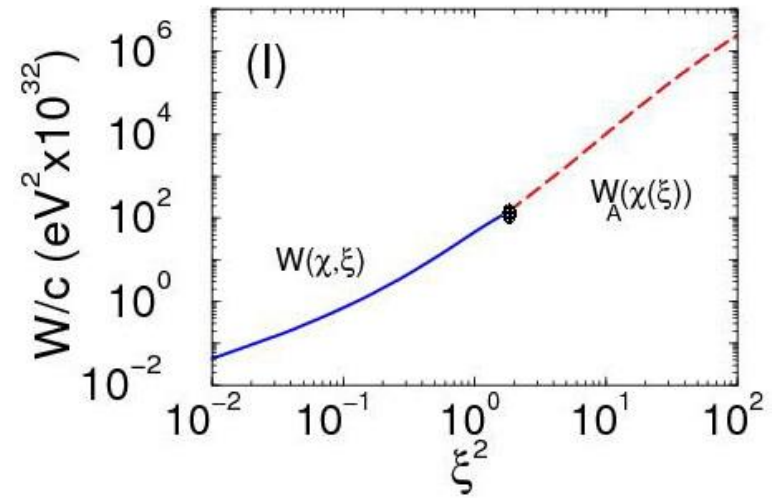
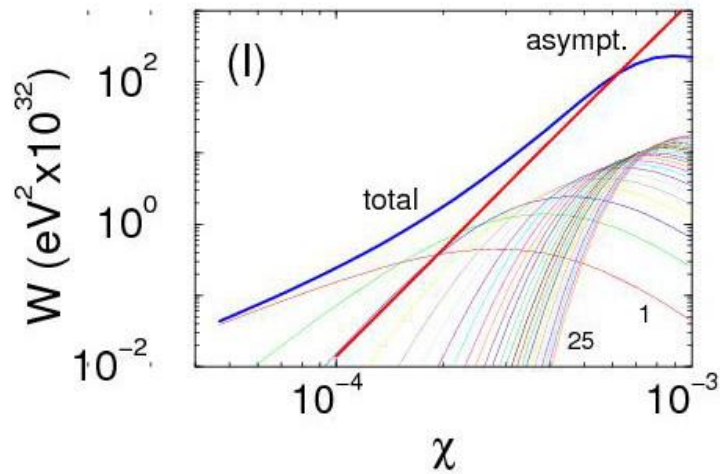
$$\sim \frac{G_F^2 m_e^6}{216\pi^3} \chi^2 \left(\ln \chi - 0.577 - \frac{1}{2} \ln 3 - \frac{5}{6} \right), \quad \chi \gg 1$$

Merenkov, Sov.J.Nucl.Phys. **42** (1985)

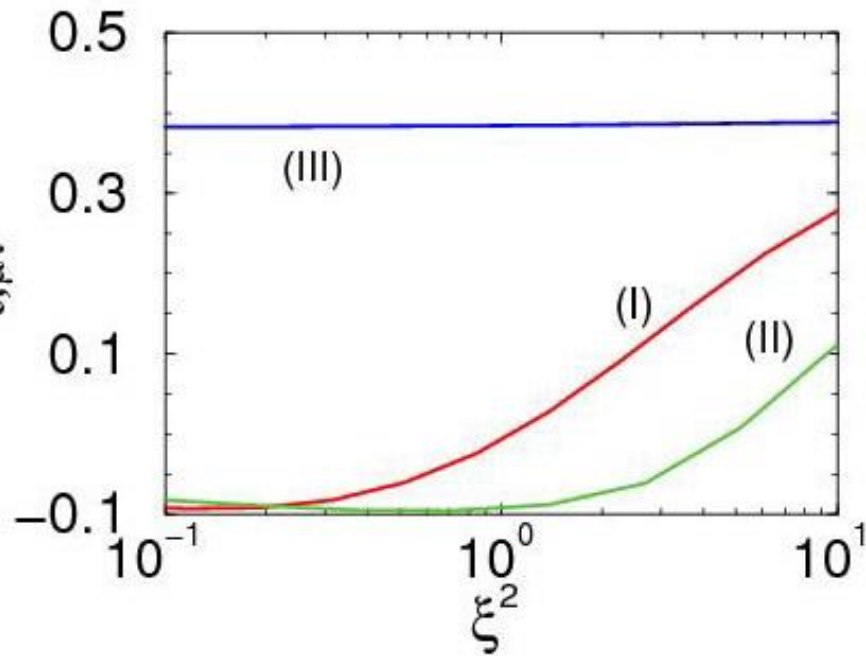
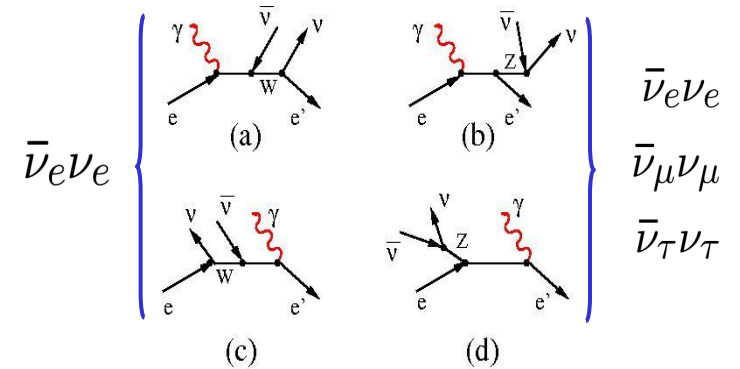
How to stitch two solutions: asymptotic and sum of finite harmonics

$$\chi^2 = \frac{\xi^2 (k \cdot p)^2}{m_e^4}$$

$$\omega_\gamma = 1.56 \text{ eV}, E_e = 40 \text{ MeV}$$



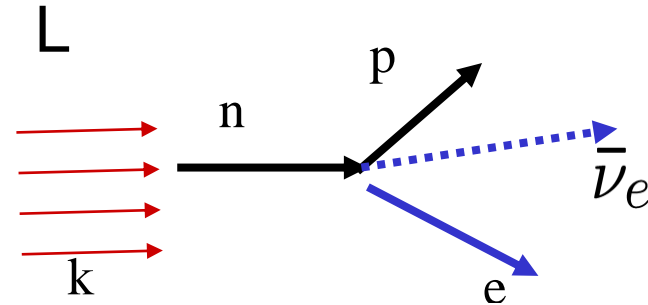
Asymmetry of production of ν_e and $\nu_\mu + \nu_\tau$



- (I) $\omega_\gamma = 1.56 \text{ eV}, E_e = 40 \text{ MeV}$
- (II) $\omega_\gamma = 1.0 \text{ KeV}, E_e = 40 \text{ MeV}$
- (III) $\omega_\gamma = 1.0 \text{ KeV}, E_e = 10 \text{ GeV}$

Neutron decay in the field of a strong electromagnetic wave

$$n \rightarrow p + e + \bar{\nu}_e$$



electron modification

$$\left\{ \begin{array}{l} \psi^D \rightarrow \psi^V \\ p \rightarrow q \\ m_e^2 \rightarrow m_*^2 = m_e^2 + e^2 a^2 = m_e^2 + m_e^2 \xi^2 \end{array} \right.$$

proton modification

$$\left\{ \begin{array}{l} p_\mu \rightarrow q_\mu = p_\mu + \frac{e^2 a^2}{2\omega_\gamma M_p} k_\mu \sim p_\mu \\ M_p \rightarrow M_{p*} = \sqrt{M_p^2 + m_e^2 \xi^2} \simeq M_p \left(1 + \frac{m_e^2}{M_p^2}\right) \simeq M_p \\ \psi^D \rightarrow \psi^V \simeq \psi^D \end{array} \right.$$

neutron modification

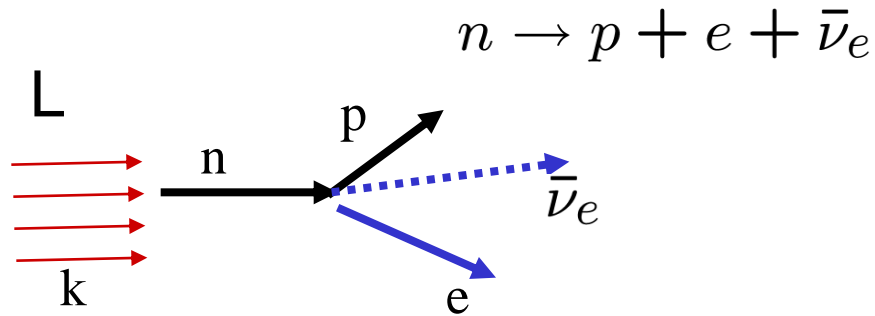
$$\left\{ \psi^D \rightarrow \psi^V \simeq \psi^D \right.$$

Amplitude

$$T_{fi} = \frac{G_F}{\sqrt{2}} \int [\bar{u}_p \gamma_\mu (1 - g_A \gamma_5) u_n] \otimes [\psi_e^* \gamma^\mu (1 - \gamma_5) v_{\bar{\nu}}] e^{-i(p_n - p_p - p_{\bar{\nu}})x} \frac{d^4 x}{\sqrt{2E_n 2E_p 2E_{\bar{\nu}}}}$$

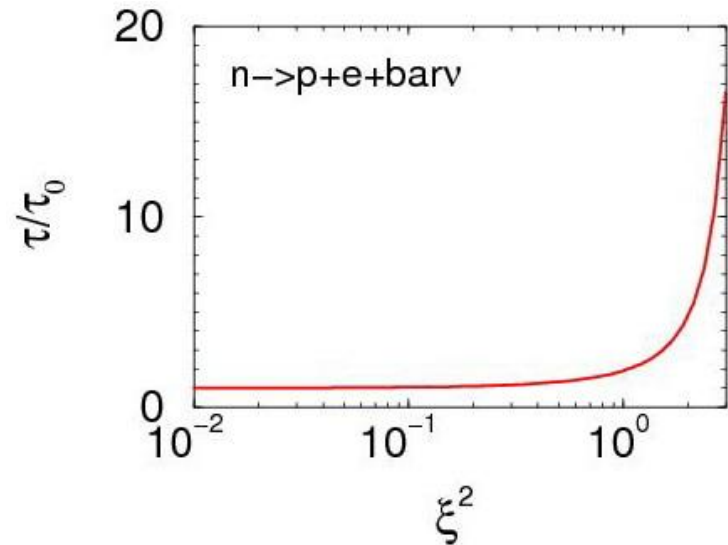
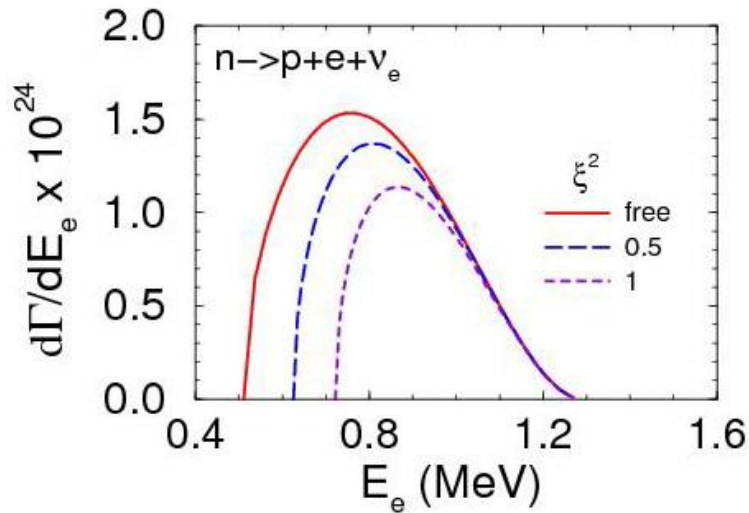
$$\left. \begin{array}{l} m_{e^*}^2 = m_e^2 + \xi^2 m_e^2 \\ m_{e^* \max} \simeq M_n - M_p \end{array} \right\} \longrightarrow \xi_{\max}^2 = \frac{(M_n - M_p)^2}{m_e^2} - 1 \simeq 5.4$$

Weak decay in the field of a strong electromagnetic wave



$$m_*^2 = m_e^2(1 + \xi^2)$$

$$\xi_{\max}^2 = \frac{(M_n - M_p)^2}{m_e^2} - 1 \simeq 5.4$$



Summary

- ★ *Strong electromagnetic fields modify basic/fundamental interactions and result in non-trivial nonlinear effects.*
- ★ *Coherent interactions with n -field photons*
- ★ *Modification of kinematical limits because (a) electron dressing and (b) coherent interactions with several photons (cumulative effect)*
- ★ *Leads to some not-trivial dynamical effects, like ν_e and $\nu_\mu + \nu_\tau$ asymmetries ...*

THE END

Thank you very much for attention !

