A METHOD OF SPALLATION NEUTRON ENERGY SPECTRUM RECONSTRUCTION WITH YTTRIUM SAMPLE ACTIVATION

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Purpose of the work

Looking for possibility to determine spallation neutron energy spectrum with $\underline{Y89}$ activation detector

Why yttrium?

Y89 the only one naturally occurring isotope

- no overlapping reactions
- easy to trace
- several residual nuclei

Resulting isotopes relatively easy to identify

Y-89(n,?) reaction cross sections available in ENDFs



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Yttrium-89 activation reactions taken into account

Reaction	Produced Isotope	T1/2	Reaction Threshold [MeV]	γ-line Energy [keV]	γ-line Intensity [%]
$V_{20}(n, q)$			6 9570*	202.51	97.3
Y89(n,g) Y90	5.1911	-0.0370	479.17	90.74	
$V_{20}(n, 2n)$	Voo	106 654	11 5	898.042	93.7
¥89(n,∠n)	100	100,050	11,5	1836.063	99.2
$V_{20}(n, 2n)$	V07	70.9h	20.9	388.53	82.00
Y89(n,3n)	107	79,011	20,0	484.805	89.7
Y89(n,4n)	Y86	14,74h	32,7	1076.64	82.00
$V_{20}(n 5n)$	VQE	2.68h	42.633	231.67	84.00
109(11,511)	COT	4.86h	42.633	231.67	22.8

Y89(n,xn) reaction residual nuclei example production. Here is, where the method idea comes from.



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Y89(n,xn) reaction residual nuclei example production – remarks



- 1. The experimental values make a curve $I(E_{thr})$
- 2. For the $I(E_{thr})$ curve the E_{thr} is a variable

The method idea

Each isotope production is a result of convoluted influence of spallation neutron spectrum and reaction cross section



The method idea

The idea is to transform one isotope production formula

$$I_{k} = Nt_{ir} \int_{E_{thr,k}}^{\infty} \varphi(E) \sigma_{k}(E, E_{thr,k}) dE$$

 $[I_{k}] - \text{nuclei/g/d}$ [N] - Y89 nuclei/g $[t_{ir}] - s$ $[\varphi(E)] - n/\text{cm}^{2}/\text{s/MeV/d}$ $[\sigma_{k}] - \text{barn} = 10^{-24} \text{cm}^{2}$

$$I_{k} \xrightarrow{\phi(E)_{t_{ir}} \to \phi(E)} N \int_{E_{thr,k}}^{\infty} \phi(E) \sigma_{k}(E, E_{thr}) dE \qquad [\phi(E)] - \text{neutron fluence/cm}^{2}/\text{MeV/d}$$

into a general one valid for any isotope

$$I(E_{thr}) = N \int_{E_{thr}}^{\infty} \phi(E) \sigma(E, E_{thr}) dE$$
 Here the E_{thr} is a variable
and solve it.

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or

The generalized equation

$$I(E_{thr}) = N \int_{E_{thr}}^{\infty} \phi(E) \sigma(E, E_{thr}) dE$$

 $\phi(E)$ looked for, neutron fluence/cm²/MeV/g/d = n/cm²/MeV/g/d $\phi(E)$ is in fact $\phi(\vec{r}, E)$ but dependence on r is not a subject $E_{thr} - a$ variable parameter $I(E_{thr})$ and $\sigma(E, E_{thr})$ to be known and continuous functions. $\phi(E) \rightarrow 0$ while $E \rightarrow \infty$

The equation creates two issues

- How to find an analytical form of $I(E_{thr})$
- Need to assume something about the $\sigma(E,E_{thr})$

 $I(E_{thr})$ case on example sample (Ed = 4.0 GeV, position P2-R3)



with κ , η and ϵ as fitting parameters

Y89(n,xn) reaction cross section case available data – experimental and evaluated



ENDF-Relational v-1.0 EXFOR Request: 6460/1, 2009-Jul-30 08:46:22

Incident Energy (MeV)

Y89(n,xn) reaction cross section case - basic assumptions and consequences

$$\sigma(E, E_{thr}) = \begin{cases} \frac{\alpha}{(kT)^{\frac{3}{2}}} \frac{(E - E_{thr})}{\sqrt{E}} \exp\left(-\frac{(E - E_{thr})}{kT}\right) & E > E_{thr} \\ 0 & E \le E_{thr} \end{cases}$$

because it fits the best EXFOR database for Y89(n,2n)Y88 reaction cross section experimental data.

The same formula applies to all Y89(n,xn) reactions. The main difference between Y89(n,2n), (n,3n) and (n,4n) reactions is their threshold energy - E_{thr} .

Notes on Y89(n,xn) reaction cross section analytical formula

- It resembles Maxwell-Boltzmann energy distribution function for ideal gas particles in equilibrium f(E) ~ sqrt(E)*exp(-E/kT).
- Statistical model of Y89 nucleus applies, i.e. nucleons behave like an ideal gas particles. The only interactions between them are elastic collisions. In an equilibrium stage their energy is dispersed around mean energy kT.
- Reaction CS proportional to f(E), σ(E)~f(E).
- N. Bohr reaction model applies to Y89(n,xn) one, where the reaction has two stages compound nucleus formation (10⁻²² s) and compound nucleus evaporation (10⁻¹⁸ 10⁻¹⁶ s).
- Y89(n,xn) reaction is caused by neutrons with energy larger than threshold energy E_{thr}.
- Y89(n,xn) reaction cross section is directly proportional to a time of flight through yttrium nucleus, i.e. inversely proportional to neutron speed, i.e. inversely proportional to the square root of neutron energy
- Neutrons and protons are the two separated, not interacting gases

Y89(n,2n)Y88 reaction cross section fitting

$$\sigma(E, E_{thr}) = \frac{\alpha}{(kT)^{\frac{3}{2}}} \frac{(E - E_{thr})}{\sqrt{E}} e^{-\frac{(E - E_{thr})}{kT}}$$

The fitting gives

- α,kT coefficients
- E_{thr} reaction threshold energy



Y89(n,xn) reaction cross section looks like

$$\sigma(E, E_{thr}) = \frac{\alpha}{(kT)^{\frac{3}{2}}} \frac{(E - E_{thr})}{\sqrt{E}} e^{-\frac{(E - E_{thr})}{kT}}$$

$$Y89(n,2n)Y88 CS fitting coefficients \alpha, kT applied to any Y89(n,xn) reaction E_{thr} - from elsewhere
Ethr - from elsewhere
$$W89(n,2n)Y88 CS fitting coefficients \alpha, kT applied to any Y89(n,xn) reaction encode the sector of the sec$$$$

Y89(n,xn) reaction cross section comparison: experimental data, MENDL-2 data base, and this work



Spallation neutron energy spectrum $\phi(E)$ determination basing on experimental data I(E_{thr}) and $\sigma(E,E_{thr})$

$$I(E_{thr}) = N \int_{E_{thr}}^{\infty} \phi(E) \sigma(E, E_{thr}) dE$$

Solving an integral equation for yttrium isotope production

Using the mentioned earlier functions for $\sigma(\text{E},\text{E}_{\text{thr}})$ and I(E_{\text{thr}}) the equation

$$I(E_{thr}) = N \int_{E_{thr}}^{\infty} \phi(E) \sigma(E, E_{thr}) dE$$

becomes

$$\frac{\alpha N}{\left(kT\right)^{\frac{3}{2}}} \int_{-\infty}^{E_{thr}} (E_{thr} - E) e^{-\frac{E}{kT}} \frac{\phi(E)}{\sqrt{E}} dE = \kappa \left(E_{thr} + \varepsilon\right) e^{-\frac{E_{thr}}{kT}}$$

Differentiating twice both sides on E_{thr} one gets:

$$\phi(E_{thr}) = \frac{\kappa(\eta + kT)^2}{\alpha N \eta^2 (kT)^{\frac{1}{2}}} \left[E_{thr} + \varepsilon - \frac{2\eta kT}{(\eta + \beta)} \right] e^{-\frac{(E_{thr} + \varepsilon)}{\eta}} \sqrt{E_{thr}}$$

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Solving a Volterra's integral equation for Yttrium isotope production - continuation

To fulfill request $\phi(E \le 0) = 0$ and $\phi(E \to \infty) \to 0$ must be

$$\varepsilon = \frac{2\eta kT}{\eta + kT}$$

and $\phi(E)$ becomes

$$\phi(E) = \frac{\kappa(\eta + kT)^2}{\alpha N\eta^2 (kT)^{\frac{1}{2}}} E^{\frac{3}{2}} e^{-\frac{1}{\eta} \left(E + \frac{2\eta kT}{\eta + kT}\right)}$$

$\phi(E)$ example view (Ed = 4 GeV, P2-R3)



The method application example

The method has been applied to "Energy plus Transmutation" experiment results elaboration

"Energy plus Transmutation" project

- International research project realised in JINR Dubna.
 - 12 states take part in
 - Started 1999
- Purpose of the project is to study transmutation on U/Pbassembly driven by accelerator NUKLOTRON.
- Transmutation samples ¹²⁹I, ²³⁷Np, ²³⁸Pu, ²³⁹Pu
- Activation detectors AI, Ti, V, Mn, Fe, Co, Ni, Cu, Y, Nb, In, Dy, Lu, W, Au, Bi
- ³He counter
- SSNTD

Yttrium samples location during irradiation





Experiments with "Energy plus Transmutation" set-up

- Proton beam:
 Deuteron beam:
- Ep = 0.7 GeV
- Ep = 1.0 GeV
- Ep = 1.5 GeV
- Ep = 2.0 GeV

- Ed = 1.6 GeV = 0.8 GeV/nucleon
- Ed = 2.52 GeV
 - = 1.26 GeV/nucleon
- Ed = 4.0 GeV
 - = 2.0 GeV/nucleon

30 Y89 samples irradiated in each experiment

Y-90 and Y-88 production spatial distribution comparison







Y-88 S2 spatial distribution based on lines 898.042 and 1836.063 keV



Ed = 1.6 GeV

Y-88 S2 spatial distribution based on lines 898.042 and 1836.063 keV



Ed = 2.52 GeV Ed = 4.0 GeV



on based on lines 898.042 and 1836.063 keV

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Y-87 and Y-86 production spatial distribution comparison











Y-86 spatial distribution based on gamma line 1076.64 keV Y86 1 OE-05 3.5E-05 5E-04 0E-04 24,0 36,2 11.8 48.4 cm Distance from the front of the Pb-target

Ed = 1.6 GeV

3 0E-05 2 5E-05 2.52 GeV 2 0E-05 euteron be 1 5E-05 0E-04 0E-0 11,8 24,0 36,2 48,4 cm Distance from the front of the Pb-target

Y-86 spatial distribution based on gamma line 1076.64 keV

Ed = 2.52 GeV Ed = 4.0 GeV



Y-85 production spatial distribution comparison







Ed = 1.6 GeV

Ed = 2.52 GeV Ed = 4.0 GeV

Function $I(E_{thr})$ in various points of R = 3 cm axis for Ed = 1.6 GeV



Ed = 1.6 GeV, axis R3 - yttrium isotope production dependence on reaction threshold energy at various axial positions

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Function $I(E_{thr})$ in various points of R = 3 cm axis for Ed = 2.52 GeV



Ed = 2.52 GeV, axis R3 - yttrium isotope production dependence on reaction threshold energy at various axial positions

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Function $I(E_{thr})$ in various points of R = 3 cm axis for Ed = 4.0 GeV



Ed = 4 GeV, axis R3 - yttrium isotope production dependence on reaction threshold energy at various axial positions.

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Ed = 1.6; 2.52 and 4.0 GeV - Neutron flux energy spectra comparison at various R3-axis positions - common Y-scale.



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Ed = 1.6; 2.52 and 4.0 GeV - Neutron flux spectra comparison at various P2 radial positions - common Y-scale



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Ed = 1.6, 2.52 and 4.0 GeV parameter kappa comparison

Ed = 1.6 GeV - Kappa



Ed = 2.52 GeV - Kappa



Ed = 4 GeV - Kappa





Ed = 1.6, 2.52 and 4.0 GeV parameter eta comparison





Ed = 2.52 GeV - eta









Ed = 1.6, 2.52 and 4.0 GeV parameter eps comparison





Ed = 2.52 GeV - eps









Ed = 1.6, 2.52 and 4.0 GeV E(ϕ_{max}) comparison

Ed = 1.6 GeV - E(fimax)



Ed = 2.52 GeV - E(fimax)



Ed = 4 GeV - E(fimax)





Ed = 1.6, 2.52 and 4.0 GeV E(ϕ_{max}) comparison



Plane P2 - Energy of spectrum maximum comparison

Energy of spectrum maximum

	P1		F	P2		P3		P4		P5	
	Ed [GeV]	Efmax	Efmax Error								
	1.6	10,13	2,33	14,17	0,16	14,75	0,36	14,58	0,26	14,27	0,36
R0	2.52			12,78	0,27						
	4.0	10,74	0,17	12,78	0,07	12,59	0,06	12,67	0,17	13,55	0,11
	1.6	10,79	0,18	12,38	0,18	12,71	0,25	13,12	0,38	14,44	0,48
R3	2.52	10,88	0,41	11,42	0,15	12,47	0,26	12,32	0,09	13,30	0,81
	4.0	11,29	0,51	12,73	0,19	12,73	0,23	12,81	0,29	13,43	0,44
	1.6			12,10	0,19						
R6	2.52			11,08	0,15						
	4.0	10,41	0,11	12,10	0,07	12,48	0,11	12,90	0,47	13,49	0,65
	1.6			11,60	0,35						
R8.5	2.52			11,01	0,25						
	4.0	10,81	0,13	11,93	0,11	11,97	0,16	12,42	0,39	13,09	0,74
	1.6			11,48	0,21						
R10.5	2.52			10,18	0,34						
	4.0	10,68	0,19	11,33	0,23	12,51	0,20	12,36	0,26	13,30	1,17
	1.6			11,30	0,21						
R13.5	2.52			10,99	0,08						
	4.0	11,00	0,15	11,45	0,16	12,34	0,51	12,49	0,51	13,07	0,68

The method error discussion

Typical errors like count statistics, deuteron fluence error, sample mass error and so on are of less importance;
Y89(n,2n)Y88 reaction cross section approximation (fitting) error of much more importance, but difficult to assess;
Y89(n,xn) reaction cross section approximation by Y89(n,2n)Y88 parameters makes additional error;
Approximation of infinite number of ideal gas particles with 50 neutrons (statistical model of nucleus);
Relativistic effects?

In some sense the method tests itself saying that $\mathcal{E} = \frac{2\eta kT}{\eta + kT}$ where both η , kT and ε are fitting parameters.

Discrepancy between the direct fitting ε value and calculated using η , kT should be an error measure of the method.

While direct fitting ϵ value ranges from 6.87 to 6.91 MeV the η , kT calculated from ranges from 7.44 to 8.5. This suggests the method error to be of order of 10-20%.

This still doesn't explain the high energy of the spectrum maximum!

Are there any other errors of the method? Note: These are the experimental values.

Yurevich's measurement of spallation neutron spectrum (but induced by protons) gave high average neutron energy too.

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What else could be done

Done

Reaction	Threshold energy [MeV]
Y89(n,g)Y90	-6,857
Y89(n,2n)Y88	11,6048
Y89(n,3n)Y87	21,0645
Y89(n,4n)Y86	33,008
Y89(n,5n)Y85	42,633

To be done

Reaction	Threshold energy [MeV]
Y89(n,He4)Rb86	-0,6916
Y89(n,p)Sr89	0,7183
Y89(n,d)Sr88	4,8995
Y89(n,2a)Br82	7,0633
Y89(n,np)Sr88	7,1494
Y89(n,T)Sr87	9,8104
Y89(n,He3)Rb87	10,0774
Y89(n,2p)Rb88	11,7317
Y89(n,2d)Rb86	23,4207
Y89(n,dT)Rb85	25,8424

But how?

Conclusions

- The method is simple, the results are coherent.
- Surprising is that using threshold detector with few threshold energies one can say so much about the spectrum in the entire energy range.
- According to the method there's no big difference between Ed = 1.6, 2.52 and 4.0 GeV spectra. The spectrum seems to have a kind of saturation at these energies of deuterons.

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Thank You for attention

Background slides

Neutron spectrum determination stages



Y89(n,2n)Y88 reaction CS EXFOR data fitting to get parameters α and kT

Any Y89(n,xn) reaction CS determination using α and kT parameters and E_{thr}

Yttrium isotope production fitting to get parameters $\kappa,\,\eta$ and ϵ

Spallation neutron spectrum determination as A function of $\alpha,$ kT, κ and $\eta.$

Y89(n,2n)Y88 reaction CS EXFOR data various fittings



Statistics	Statistics
Reduced Chi-Sqr	Adj. R- Square
0,00929	0,88299

Statistics	Statistics
Reduced Chi-Sqr	Adj. R- Square
0,0112	0,85882

Statistics		Statistics	3	
Reduced Chi-Sqr		Adj. R- Square		
0,01045		0,86833		
n	S	tatistics	Stati	stics
Value	R C	educed hi-Sqr	Adj. R- Square	
0,84421	0.	,0085	0,89	293

Y89(n,2n)Y88 reaction CS EXFOR data various fittiings

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Y89(n,2n)Y88 CS EXFOR data various fittings – cont.

	alfa	alfa	kT	kT	Ethr	Ethr	Statistics	Statistics
	Value	Standard	Value	Standard	Value	Standard	Reduced Chi-	Adj. R-
		Error		Error		Error	Sqr	Square
Y89(n,2n)Y88	41,71254	1,16615	8,02267	0,26505	11,84393	0,05482	0,00929	0,88299
CS								

	alfa	alfa	kT	kT	Ethr	Ethr	Statistics	Statistics
	Value	Standard	Value	Standard	Value	Standard	Reduced Chi-	Adj. R-
		Error		Error		Error	Sqr	Square
Y89(n,2n)Y88	21,68922	0,52543	4,33228	0,13764	10,59808	0,11516	0,0112	0,85882
CS								

	alfa	alfa	kT	kT	Ethr	Ethr	Statistics	Statistics
	Value	Standard	Value	Standard	Value	Standard	Reduced Chi-	Adj. R-
		Error		Error		Error	Sqr	Square
Y89(n,2n)Y88	52,94533	2,71832	18,60426	1,20092	12,5678	0,03074	0,01045	0,86833
CS								

	alfa	alfa	kТ	kT	Ethr	Ethr	n	n	Statistics	Statistics
	Value	Standard	Value	Standard	Value	Standard	Value	Standard	Reduced	Adj. R-
		Error		Error		Error		Error	Chi-Sqr	Square
Y89(n,2n)Y88	43,45287	1,90045	9,11047	0,96073	12,16393	0,08633	0,84421	0,06974	0,0085	0,89293
CS										

Evaporation mechanism of Y89(n,xn) reaction

○ → Incident neutron energy

