

Relativistic Mean-Field Model with Scaled Hadron Masses and Couplings.

A.Khvorostukhin in collaboration with
V.Toneev and D.Voskresensky

- Lagrangian, pressure and equations of motion
- Determination of parameters
- Model properties at $T=0$ and $T \neq 0$
- Application to HIC
- Conclusions and perspectives

Introduction: Model choice

EoS for hot and dense matter

A.Khvorostukhn et al.

Lagrangian, pressure and equations of motion

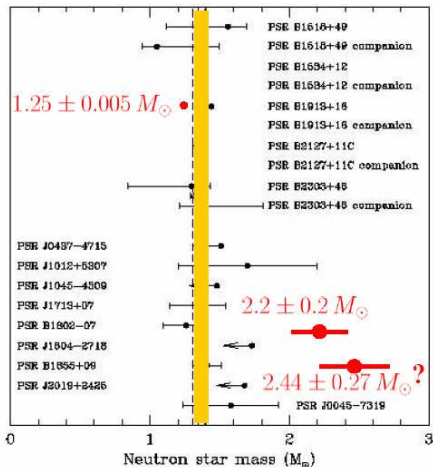
Determination of parameters

Model properties

Application to HIC

Conclusions and perspectives

Neutron star masses



- EoS of Urbana–Argonne group (A18+ δv +UIX*)
 - There might exist heavy NSs
- ⇒ **KVOR**

E.Kolomeitsev, D.Voskresensky, Nucl. Phys. **A759** (2005) 373

We want to generalize the RMF model with density-dependent hadron masses and couplings to finite temperatures for further application to hydrodynamics. (Nucl. Phys. **A 791** (2007),180, nucl-th/0802.3999)

- mean-field thermodynamics for density-dependent model
- lowest baryon resonances and Goldstone bosons
- excitations of σ , ω and ρ fields and their interactions (beyond MF)
- properties of dense and hot hadronic matter

Lagrangian

EoS for hot and dense matter

A.Khvorostukhn et al.

Lagrangian, pressure and equations of motion

Determination of parameters

Model properties

Application to HIC

Conclusions and perspectives

We start from generalized KVOR based lagrangian and **expand it retaining only quadratic terms in the fields of excitations.**

$$\begin{aligned} \widetilde{\mathcal{L}}[f, \omega_0, R_0] = & \underbrace{\sum_b \Psi_b^+ (i\gamma^0 \gamma^\mu \partial_\mu - X_b - \beta m_b^*) \Psi_b}_{\text{baryons}} \underbrace{- \widetilde{U}(f) + V(f)}_{\mathcal{L}_{MF}} \\ & + \underbrace{\mathcal{L}_{\sigma'} + \mathcal{L}_{\omega'} + \mathcal{L}_{\rho'}}_{\text{ideal gases of excitations}} + \underbrace{\mathcal{L}_{\pi^0} + \mathcal{L}_{\eta'} + \mathcal{L}_\phi + \mathcal{L}_{K^*}}_{\text{ideal gases}} + \mathcal{L}_\eta + \sum_{\varphi \in G} \mathcal{L}_\varphi \end{aligned}$$

$$\mathcal{L}_\varphi = [(\partial_\mu - iX_\phi \delta_{\mu 0}) \varphi^+] (\partial^\mu + iX_\phi \delta^{\mu 0}) \varphi - m_\varphi^{*2} \varphi^+ \varphi$$

where mean field approximation is already taken into account:

$$\begin{aligned} X_b &= g_{\omega b} \chi_\omega \omega_0 + g_{\rho b} t_b^3 \chi_\rho R_0, & X_\varphi &= g_{\omega \varphi}^* \omega_0 + g_{\rho \varphi}^* t_\varphi^3 R_0 \\ V(f) &= \frac{1}{2} m_\omega^{*2} \omega_0^2 + \frac{1}{2} m_\rho^{*2} R_0^2 \\ \widetilde{U}(f) &= \frac{1}{2} m_\sigma^{*2} \sigma^2 + U(f) = \frac{m_N^4 f^2}{2C_\sigma^2} \eta_\sigma(f) + U(f) \end{aligned}$$

$$\text{with } m_N^* = m_N \Phi_N(f), \quad m_b^* = m_b - g_{\sigma B} \chi_{\sigma} \sigma,$$

$$m_j^* = m_j \Phi_j(f), \quad m_{\varphi}^* = m_{\varphi} - g_{\sigma\varphi}^* \sigma,$$

and scaling functions $\chi_j = \chi_j(f)$, $j \in \mathcal{M}$

Non-linear self-interaction term

$$U = m_N^4 \left(\frac{b}{3} f^3 + \frac{c}{4} f^4 \right)$$

The variable has been introduced

$$f = \frac{g_{\sigma N} \chi_{\sigma} \sigma}{m_N}$$

$$C_i = \frac{g_{iN} m_N}{m_i}, \quad \eta_i(f) = \frac{\Phi_i^2(f)}{\chi_i^2(f)}, \quad g_{iH} = x_{iH} g_{iN}, \quad i = \sigma, \omega, \rho$$

$$\text{Sets of hadrons } B = \{N, \Delta, \Lambda, \Sigma, \Xi, \Sigma^*(1385), \Xi^*(1530), \Omega^-\}$$

$$G = \{\pi^+, K\}, \quad \mathcal{M} = \{\sigma, \omega, \rho\}$$

Pressure and equations of motions

Then the pressure is given by

$$P[f, \omega_0, R_0] = \sum_b (p_b + \bar{p}_b) + \sum_\phi (p_\phi + \bar{p}_\phi) + \sum_{j \in \mathcal{M}} p_j \\ + p_{\pi^0} + p_\eta + p_{\eta'} + p_\phi + p_{K^*} + \underbrace{V(f) - \tilde{U}(f)}_{P_{MF}[f, \omega_0, R_0]}$$

where p_j – the pressure of ideal gas of particles (\bar{p}_j – antiparticles) with mass m^* and effective chemical potential

$$\mu_j^* = \mu_j - X_j = b_j \mu_B + s_j \mu_S - X_j$$

Self-consistency conditions

$$\frac{\partial P}{\partial \omega_0} = 0, \quad \frac{\partial P}{\partial R_0} = 0, \quad \frac{\partial P}{\partial f} = 0$$

are equivalent to equations of motions

Masses of excitations

$$m_{\omega'}^{*2} = \frac{\partial^2}{\partial \omega_0^2} P_{MF}[f, \omega_0, R_0] = m_{\omega}^{*2}$$

$$m_{\rho'}^{*2} = \frac{\partial^2}{\partial R_0^2} P_{MF}[f, \omega_0, R_0] = m_{\rho}^{*2},$$

$$m_{\sigma'}^{*2} = -\frac{\partial^2}{\partial \sigma^2} P_{MF}[f, \omega_0(f), R_0(f)] = -\left(\frac{df}{d\sigma}\right)^2 \frac{d^2 P_{MF}}{df^2}$$

Why is P_{MF} used in these formulaes?

- We keep only quadratic terms in lagrangian, so $P_{bos.es}$ is disregarded.
- Including P_B is equal to taking into account of the baryon loops that leads to some problems. Our aim is the construction of a thermodynamically consistent model so in the final version of our model we suppress P_B to keep thermodynamical consistency and for simplicity sake.

Basic parameters

We take the Brown-Rho scaling ansatz (1991) in the simplest form

$$\Phi_N = \Phi_\sigma = \Phi_\omega = \Phi_\rho = 1 - f \quad \left(\Phi = \frac{m^*}{m} \right)$$

An appropriate behavior of EoS for $T = 0$ is obtained with

$$\eta_\sigma = 1, \quad \eta_\omega = \frac{1 + zf_0}{1 + zf}, \quad \eta_\rho = \frac{\eta_\omega}{\eta_\omega + 4 \frac{C_\omega^2}{C_\rho^2} (\eta_\omega - 1)}$$

with $f_0 \equiv f(T = 0, n_0)$, $z = 0.65$

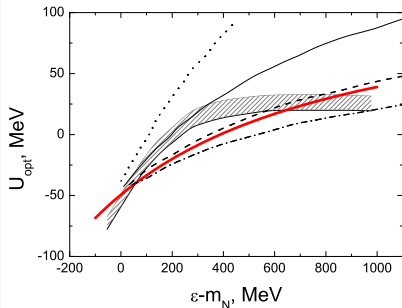
Input parameters (at the saturation point)

$$n_0 = 0.16 \text{ fm}^{-3}, \quad E_{bind} = -16 \text{ MeV}, \quad K = 275 \text{ MeV}$$

$$m_N^*/m_N = 0.805$$

- $x_{\omega H}$ – from quark counting
- free pion

Nucleon optical potential

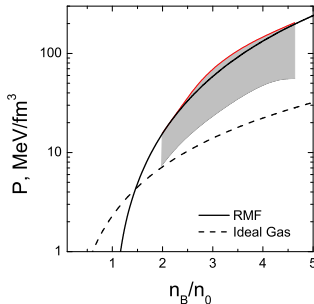


$m_N^*/m_N = 0.54, 0.71, 0.8, 0.85$ (dotted, thin continuous, dashed and dashed-dotted lines)

Shaded area - uncertainties in extrapolation from finite nuclei to cold nuclear matter

H.Feldmeier, J. Lindner, Z.Phys. A341 (1991) 83

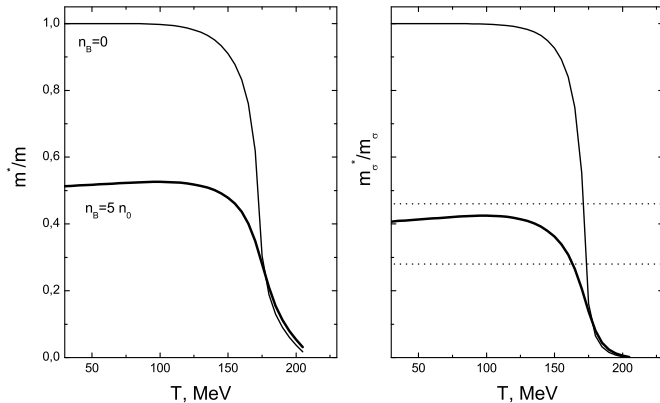
Pressure



P. Danielewicz, nucl-th/0512009

Effective masses

$T \neq 0$. Effective-to-bare mass ratios



N, ω, ρ masses

σ mass

dotted straight lines show the σ mass levels $m_\sigma = 2m_\pi$ and m_π

$T \neq 0$. Kaon energies

EoS for hot and dense matter

A.Khvorostukhn et al.

Lagrangian, pressure and equations of motion

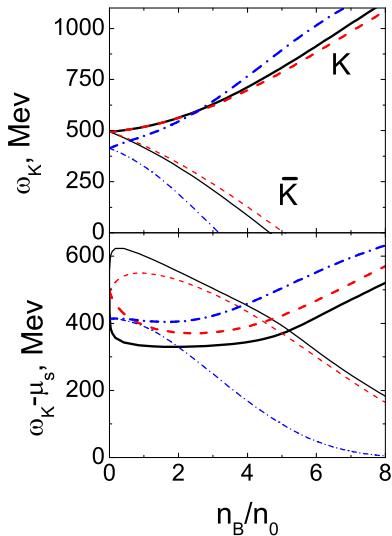
Determination of parameters

Model properties

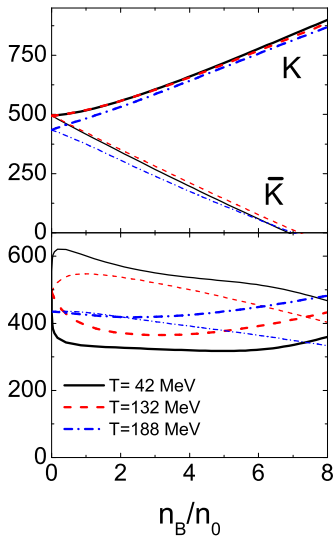
Application to HIC

Conclusions and perspectives

Left panel: set (A) without scaling



Right panel: set (B) with scaling



T-dependence of pressure and specific heat

EoS for hot and dense matter

A.Khvorostukhn et al.

Lagrangian, pressure and equations of motion

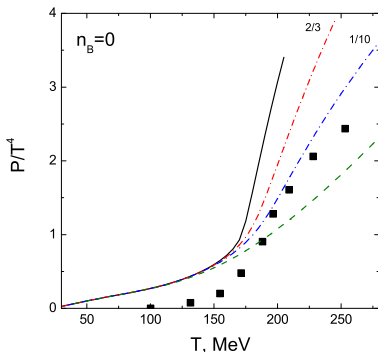
Determination of parameters

Model properties

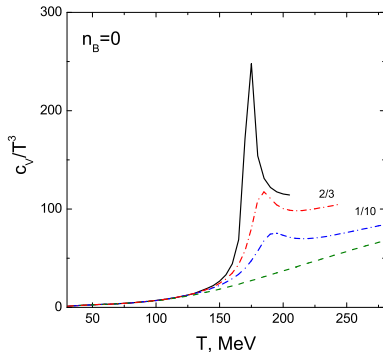
Application to HIC

Conclusions and perspectives

Pressure



Specific heat



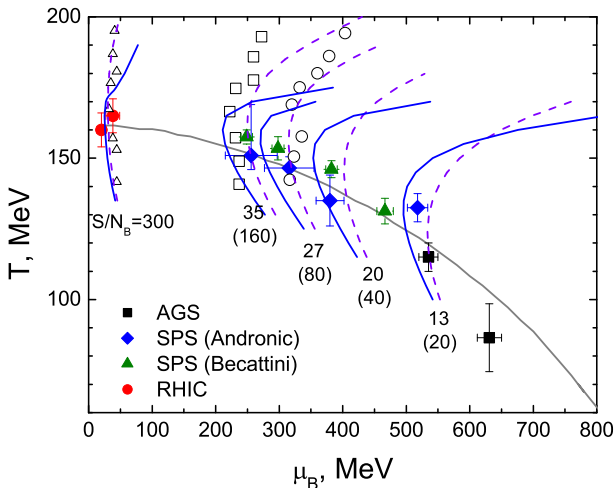
dashed - ideal gas

solid - our model

dashed-dotted, dashed-dotted - our model with suppressed σ -couplings

squares - lattice QCD for the 2+1 flavor

Isentropic trajectories for RHIC and SPS



Our S/N_B from hydro-calculations by
M. Reiter, A. Dumitru, J. Brachman, J.A. Maruhn, H. Stöcker and W. Greiner, nucl-th/9801068

M. Reiter, A. Dumitru, J. Brachman, J.A. Maruhn, H. Stöcker and W. Greiner, nucl-th/9801068

Results and perspectives

EoS for hot and dense matter

A.Khvorostukhn et al.

Lagrangian, pressure and equations of motion

Determination of parameters

Model properties

Application to HIC

Conclusions and perspectives

- The modified RMF $\sigma - \omega - \rho$ model with scaled hadron masses and couplings (SHMC model) was generalized to finite temperatures, simulating a chiral symmetry restoration with a temperature increase.
- Besides nucleon and mean fields, the model includes low-lying baryon resonances and their antiparticles, boson excitations (Goldstone bosons) and σ -, ω -, ρ -excitations.
- The EoS for $T=0$ satisfies general constraints known from atomic nuclei, neutron stars and those coming from the flow analysis of HIC data.
- Study of thermodynamic properties of hot and dense hadronic matter was started with the constructed EoS.
- We want to extend the model by matching it with a quark phase, including finite-size effects of a mixed phase and applying to hydrodynamics (dynamics and observables)

EoS for hot and dense matter

A.Khvorostukhn
et al.

Lagrangian,
pressure and
equations of
motion

Determination
of parameters

Model
properties

Application to
HIC

Conclusions and
perspectives

Parameters of hyperons, kaons and pions

EoS for hot and dense matter

A.Khvorostukhn et al.

Lagrangian, pressure and equations of motion

Determination of parameters

Model properties

Application to HIC

Conclusions and perspectives

From quark counting

$$x_{\omega\Lambda} = x_{\omega\Sigma} = x_{\omega\Sigma^*} = 2x_{\omega\Xi} = x_{\rho\Lambda} = x_{\rho\Sigma} = x_{\rho\Sigma^*} = 2x_{\rho\Xi} = \frac{2}{3}$$

$$x_{\omega K} = \frac{1}{3}$$

$x_{\sigma b}$ for Λ , Σ , Ξ are found from

$$E_{bind}^{hyp} = \mu_{hyp}(T=0, n_0) - m_{hyp} = \frac{C_\omega^2}{m_N^2 \eta_\omega(f_0)} x_{\omega B} n_0 - x_{\sigma B} (m_N - m_N^*(n_0))$$

with $E_{bind}^\Lambda = -30 \text{ MeV}$, $E_{bind}^\Sigma = -10 \text{ MeV}$, $E_{bind}^\Xi = -18 \text{ MeV}$
 $x_{\sigma K}$ is deduced from

$$U_{K^-}(T=0, n_0) = -g_{\sigma K}^* \sigma - g_{\omega K}^* \omega_0$$

with $U_{K^-}(n_0) = -120 \div -130 \text{ MeV}$, $U_{K^+}(n_0) = 20 \div 30 \text{ MeV}$

$$(A): \quad g_{\omega K}^* = x_{\omega K} g_{\omega N}, \quad g_{\sigma K}^* = x_{\sigma K} g_{\sigma N}$$

$$(B): \quad g_{\omega K}^* = x_{\omega K} g_{\omega N} \frac{\chi_\omega}{\chi_\omega(f_0)}, \quad g_{\sigma K}^* = x_{\sigma K} g_{\sigma N} \frac{\chi_\sigma}{\chi_\sigma(f_0)}$$

$$x_{\omega\pi} = x_{\sigma\pi} = 0 \quad (\text{free pion})$$

Δ -baryon. Nonzero width

EoS for hot and dense matter

A.Khvorostukhn et al.

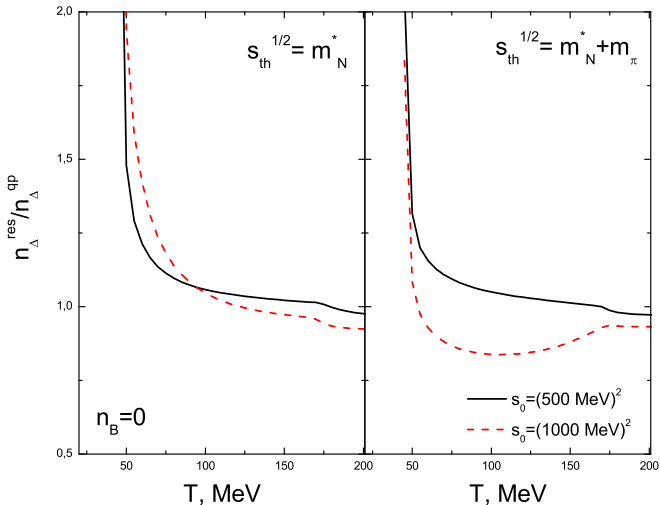
Lagrangian, pressure and equations of motion

Determination of parameters

Model properties

Application to HIC

Conclusions and perspectives



ρ -meson. Nonzero width

EoS for hot and dense matter

A.Khvorostukhn et al.

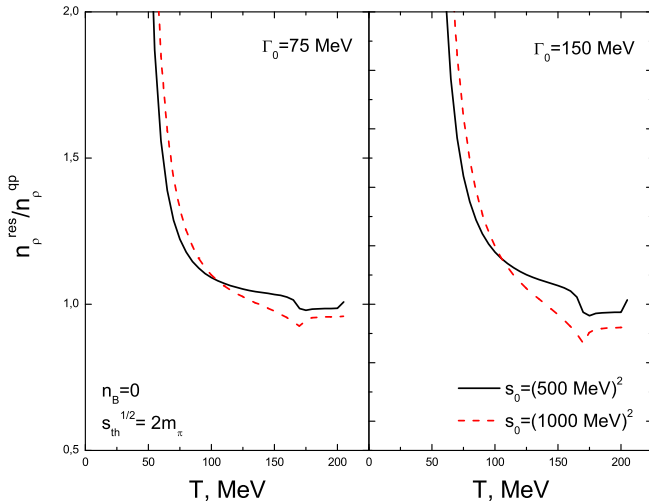
Lagrangian, pressure and equations of motion

Determination of parameters

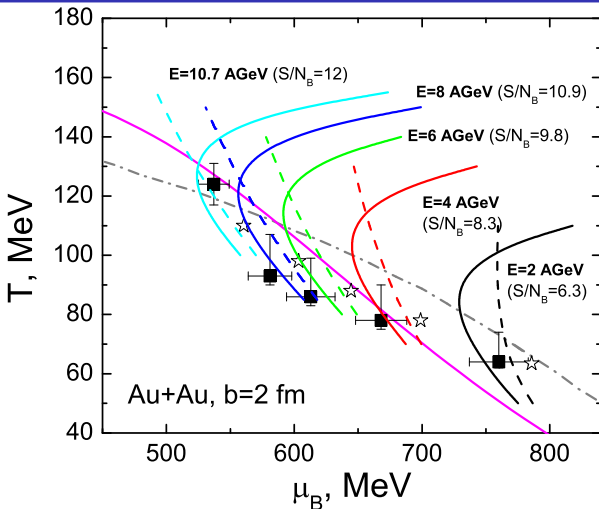
Model properties

Application to HIC

Conclusions and perspectives



Isentropic trajectories for AGS



Solid lines - our model, dashed lines - ideal gas model
thin solid, dash-dotted lines - freeze-out curves
squares - experimental points
 S/N_B from QGSM model

A. Andronic, P. Braun-Munzinger and J. Stachel, Nucl. Phys. **A772** (2006) 167

Particle ratios for other energies

EoS for hot and dense matter

A.Khvorostukhn et al.

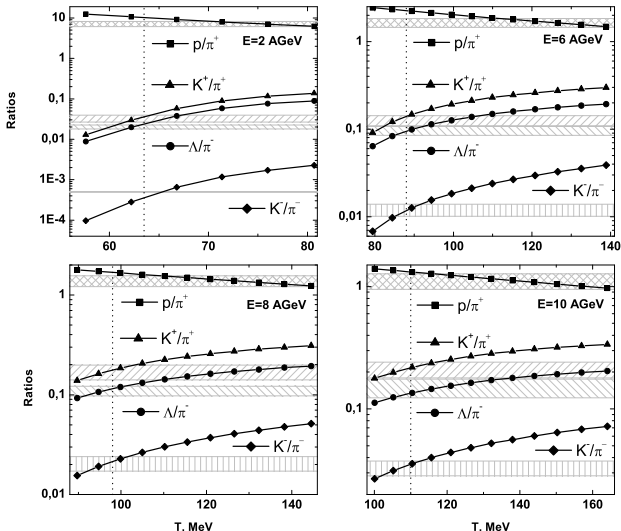
Lagrangian, pressure and equations of motion

Determination of parameters

Model properties

Application to HIC

Conclusions and perspectives



Strangeness suppression was taken into account!
Shaded bands correspond to uncertainties of experimental data

A. Andronic, P. Braun-Munzinger and J. Stachel, Nucl. Phys. **A772** (2006) 167